

The background of the image is a deep space field filled with numerous galaxies of various colors (yellow, orange, blue, purple) and sizes. Overlaid on this field is a network of thin, bright blue lines that represent the cosmic web, connecting different regions of the universe. The text 'Just Fibbers' is centered in the lower half of the image. 'Just' is in a smaller, bold, white sans-serif font, while 'Fibbers' is in a much larger, bold, white sans-serif font.

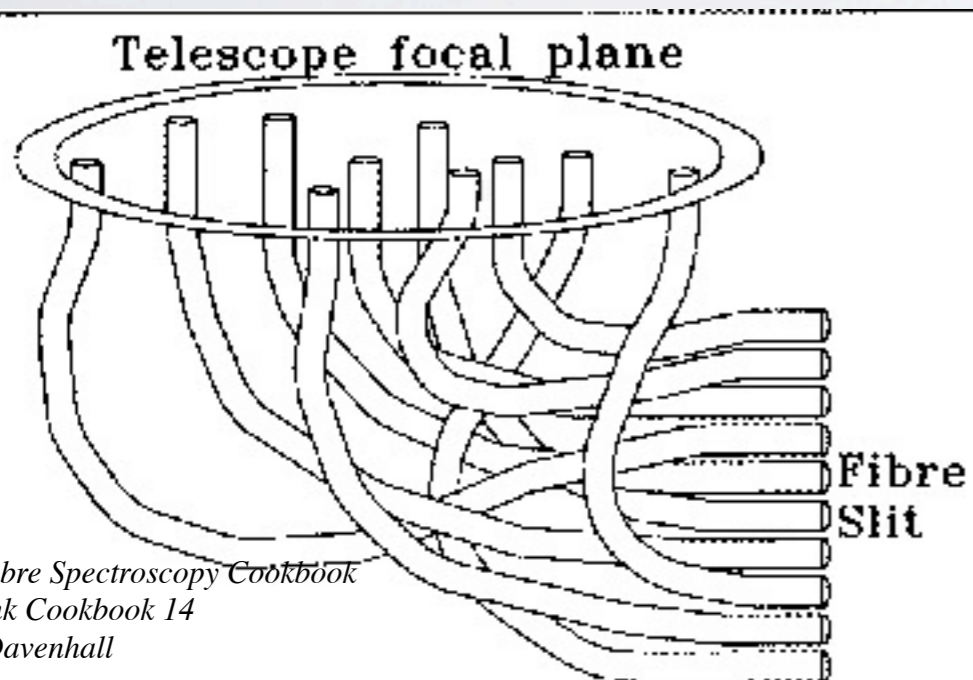
**Just  
Fibbers**

# OPTICAL FIBER TESTING AT IEU

- This lecture is about my activities testing optical fibers for the BigBOSS (now known as MS-DESI) experiment, in the lab at the Institute for the Early Universe, Seoul.
- I hope you have an opportunity to visit a lab to see how optics measurements are performed.

# FIBER SPECTROGRAPHS

- “Longslit” spectrographs are used to make one spectrum per pointing of telescope.
  - NO GOOD for large statistical studies or surveys with many objects
- Fiber-Feed spectrographs “multiplex” - you put a fiber at the position of each galaxy, fibers carry light to spectrograph

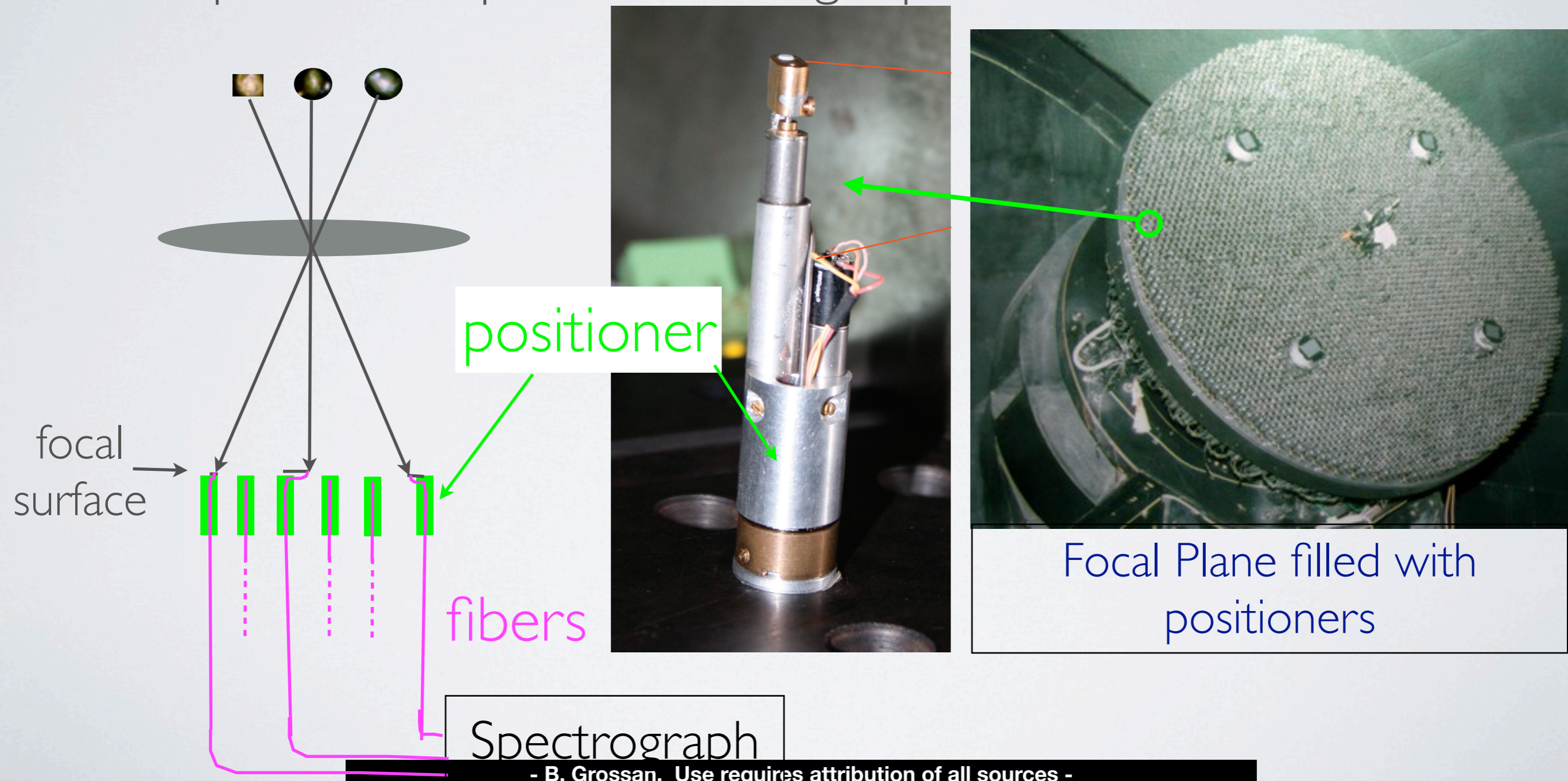


to spectrograph

**Use requires attribution of all sources -**

# KEY TO DESI INSTRUMENT: 5000 SPECTRA PER POINTING VIA ROBOT POSITIONERS

- 5000 positioners put fiber at target position



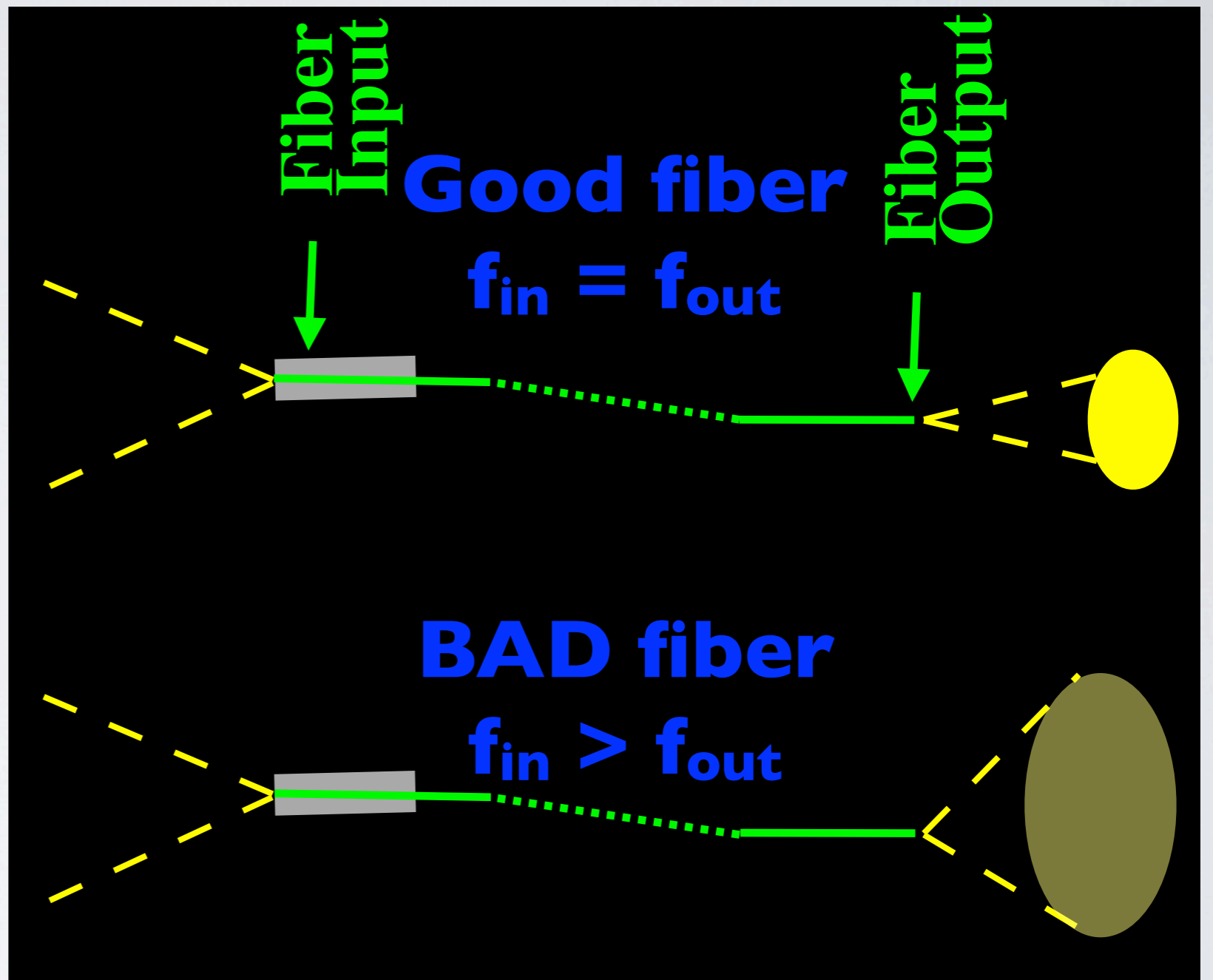
# WHY TEST OPTICAL FIBERS?

## ■ Focal Ratio

### Degradation =

when a fiber spreads out the incoming beam too much.

- Beam spread is BAD in astronomy - multiplies read noise & background
- BOSS experiment had many runs of fibers w/ bad FRD - must test.



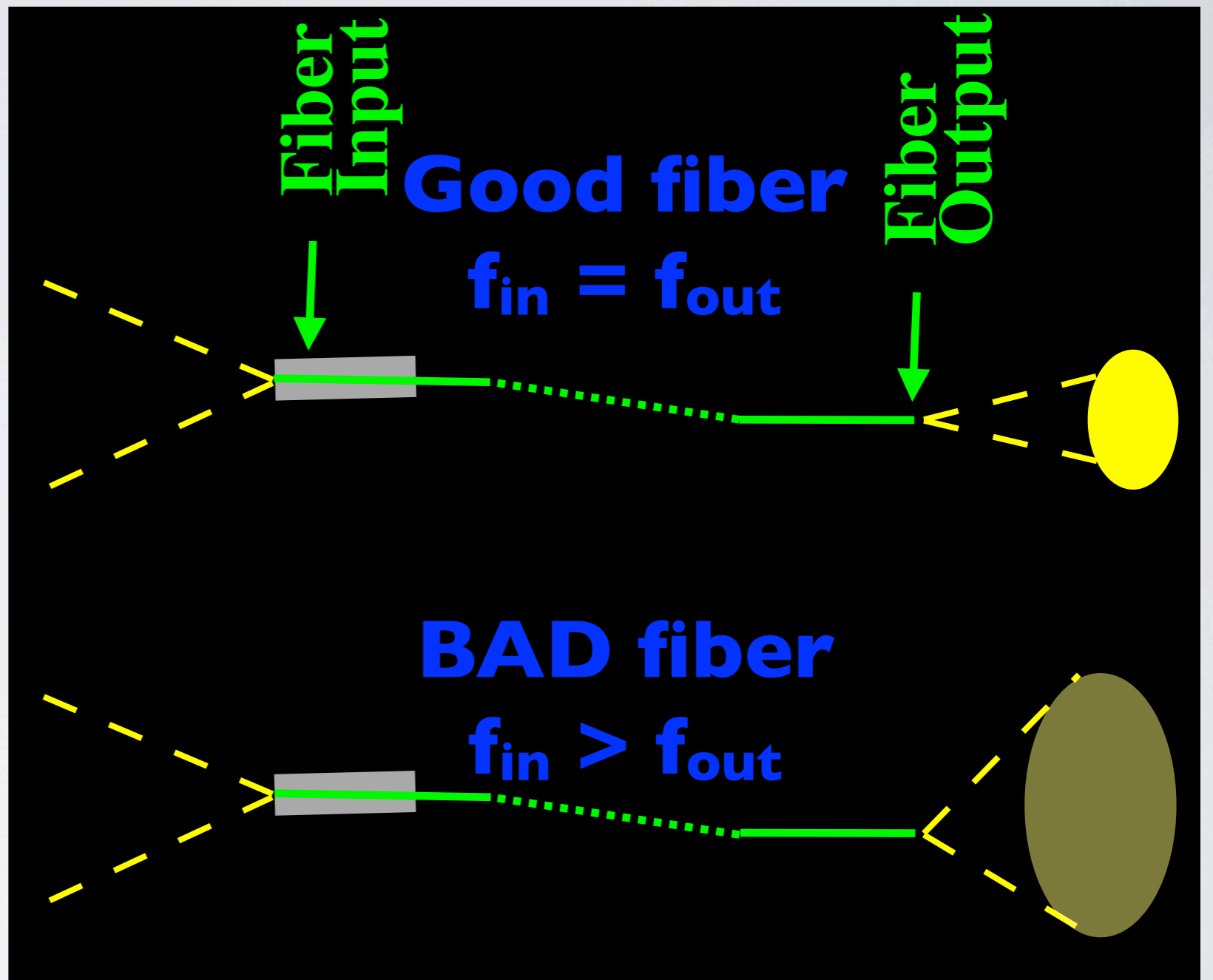
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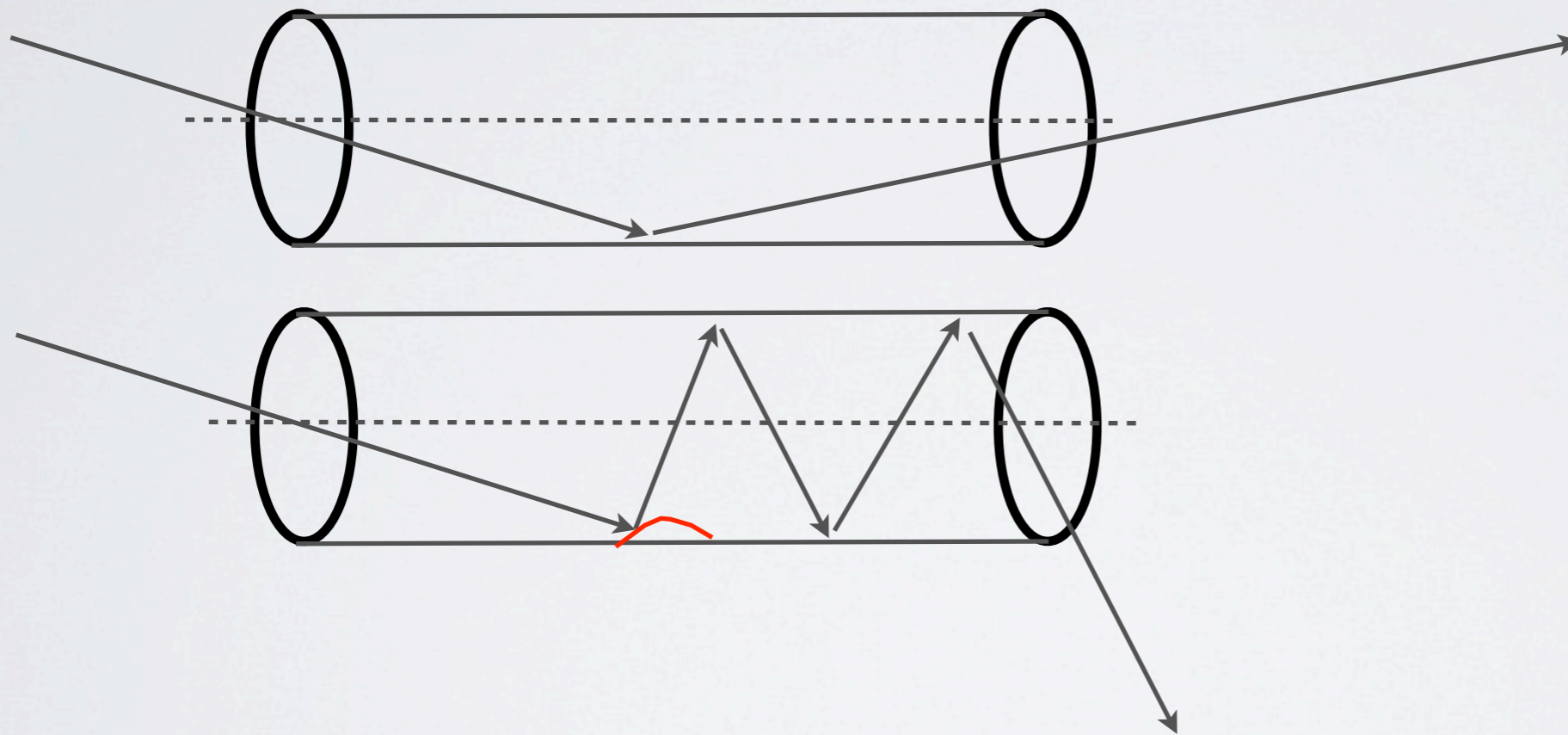
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Critical- **all 5000** fibers must be tested for FRD!

# WHAT CAUSES FRD?

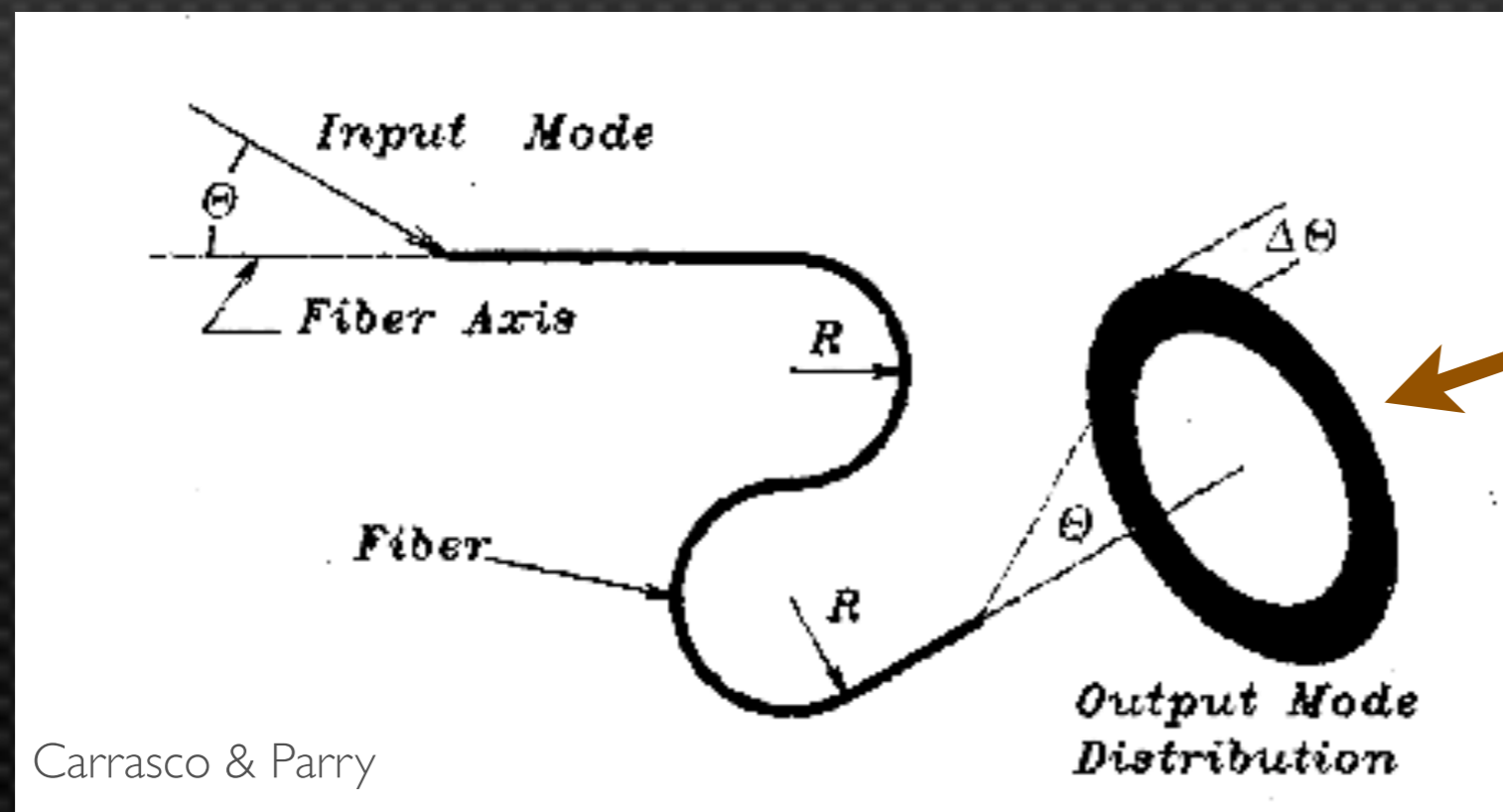
- Theory says microbends cause scattering of rays off of ideal path.



- the smaller the characteristic microbend radius, the worse the performance of the fiber. We measure this with the laser.

# Measure-Collimated Light Test

- Shine LASER or other collimated source in fiber, look at ring pattern out
  - Width of ring must reflect variation of ray path through fiber
  - Can be translated into FRD measure



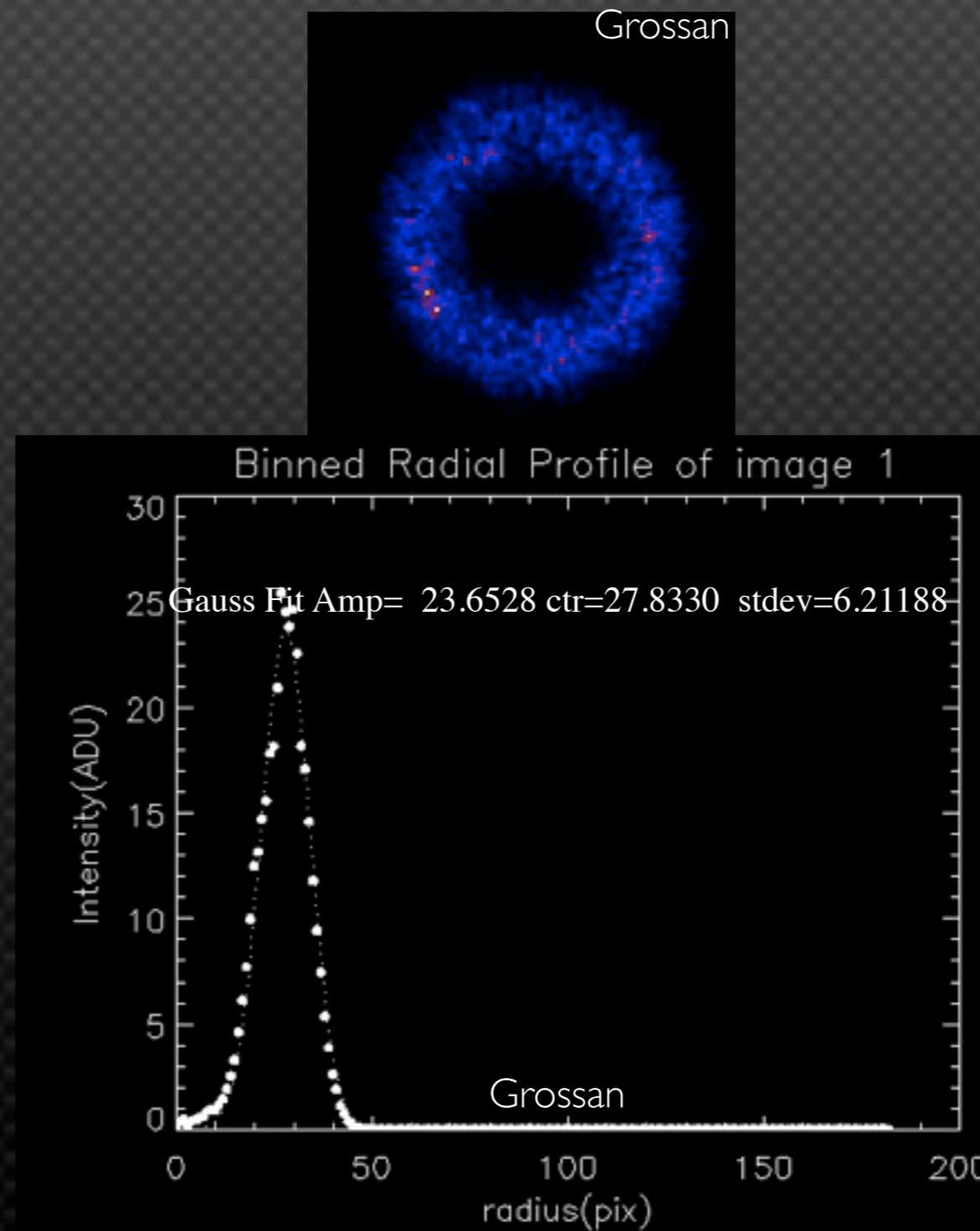
**Narrow ring=good**

**Fat ring = bad**



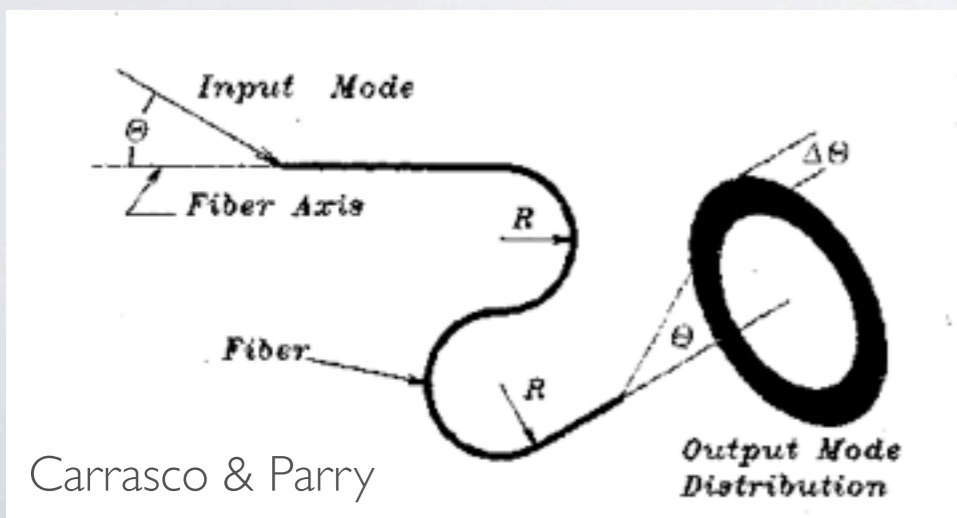
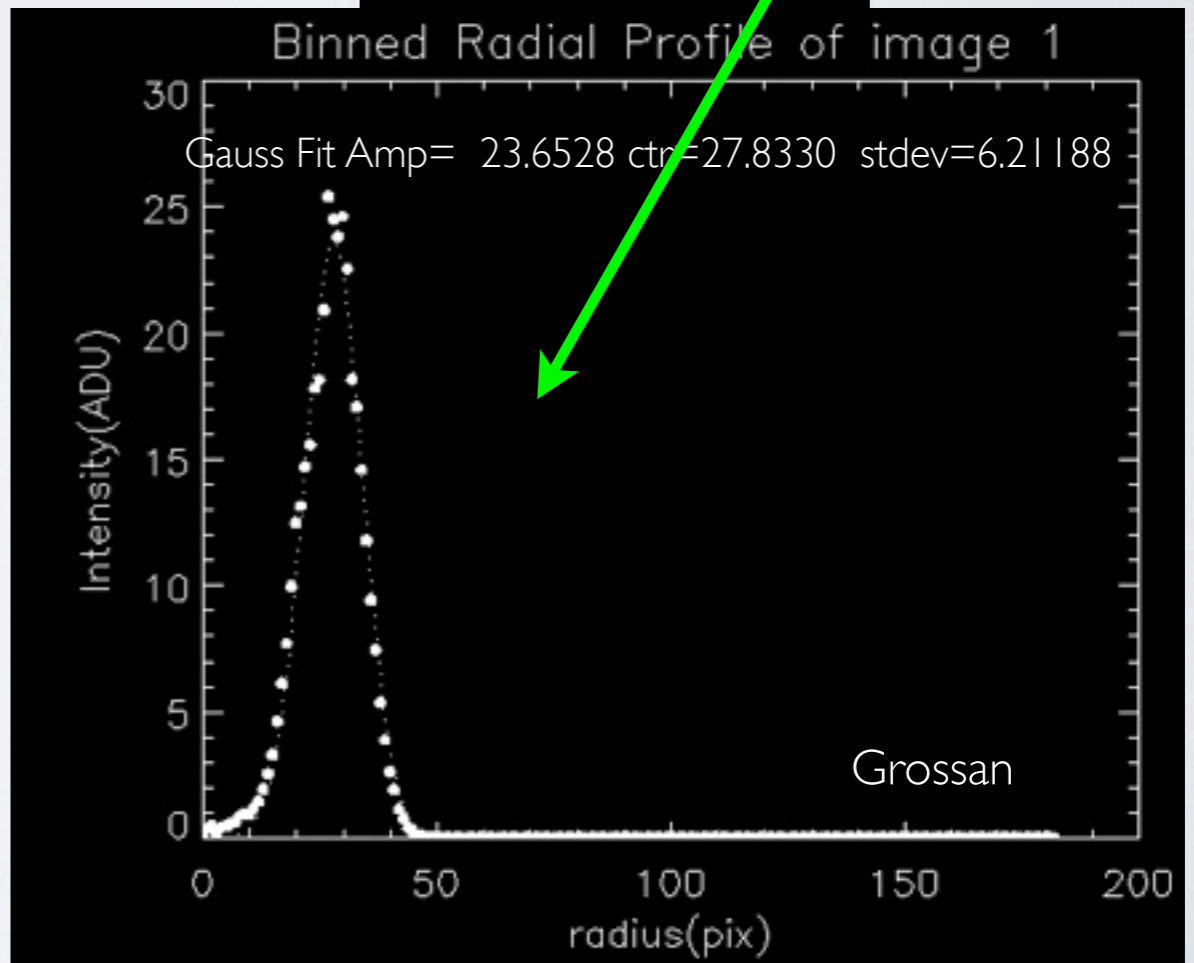
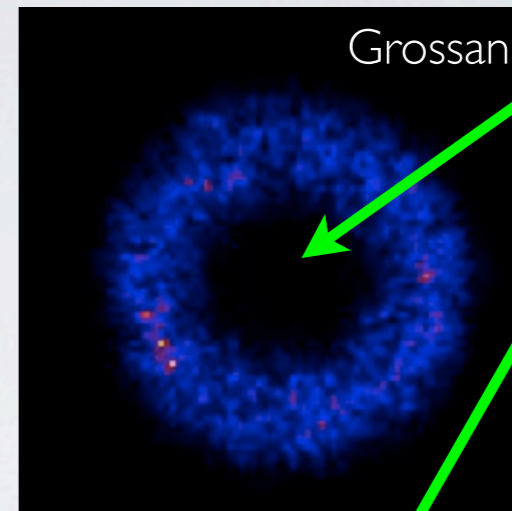
# Measurements

- Laser "Rings" AKA coherent Source Annulus

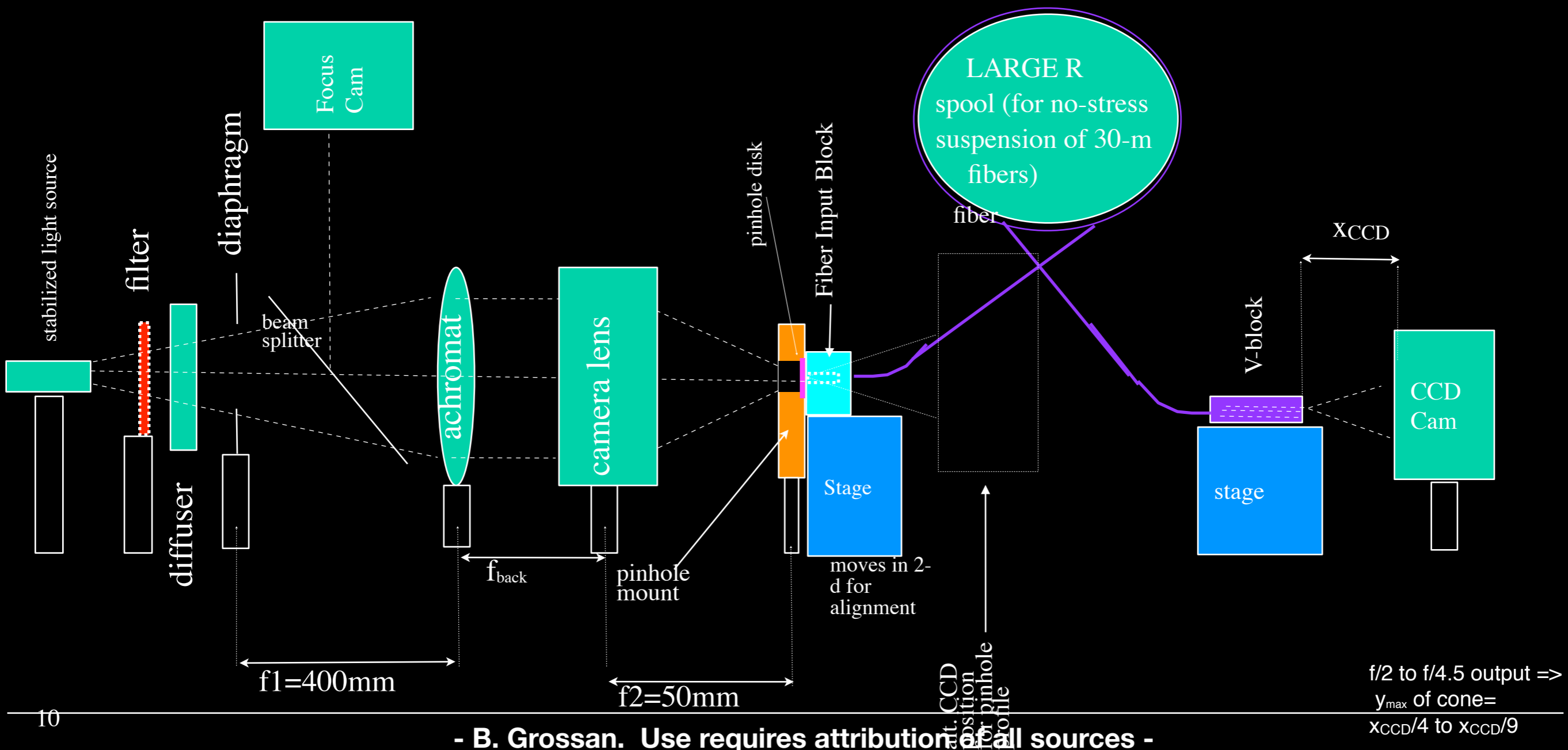


# EXERCISE

- How do you reduce these data to obtain this?
- How do you reduce these data to tell me the angular spread of the annulus (bagel)?
  - Hint: Draw the experiment, think about what you need to measure.
- Write the complete algorithm and program. Data are coming.



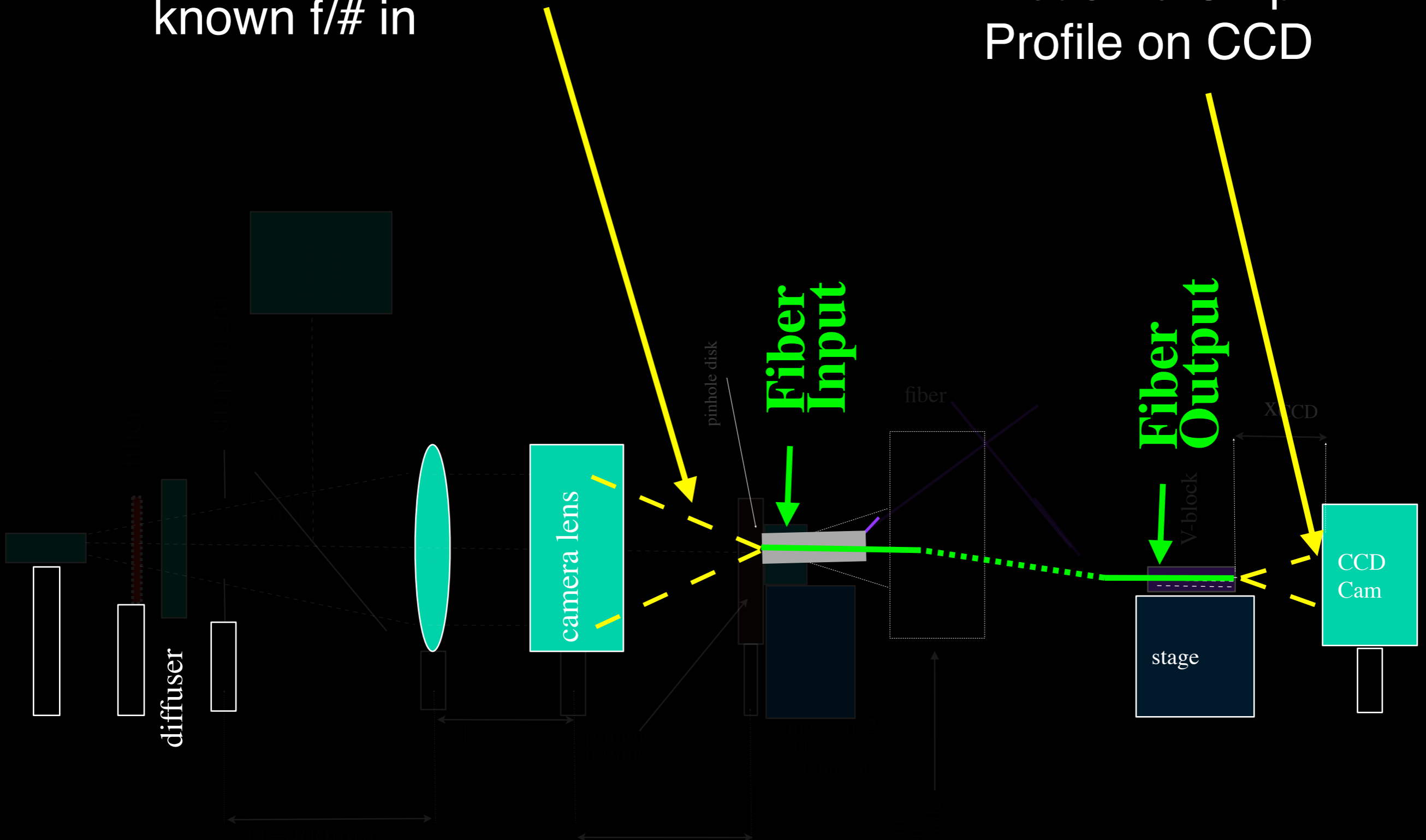
# IBBOSS Lamp Setup (after Carrasco & Parry 94)



# Simplified- Incoherent Beam Test

- Shine well-defined beam of known f/# in

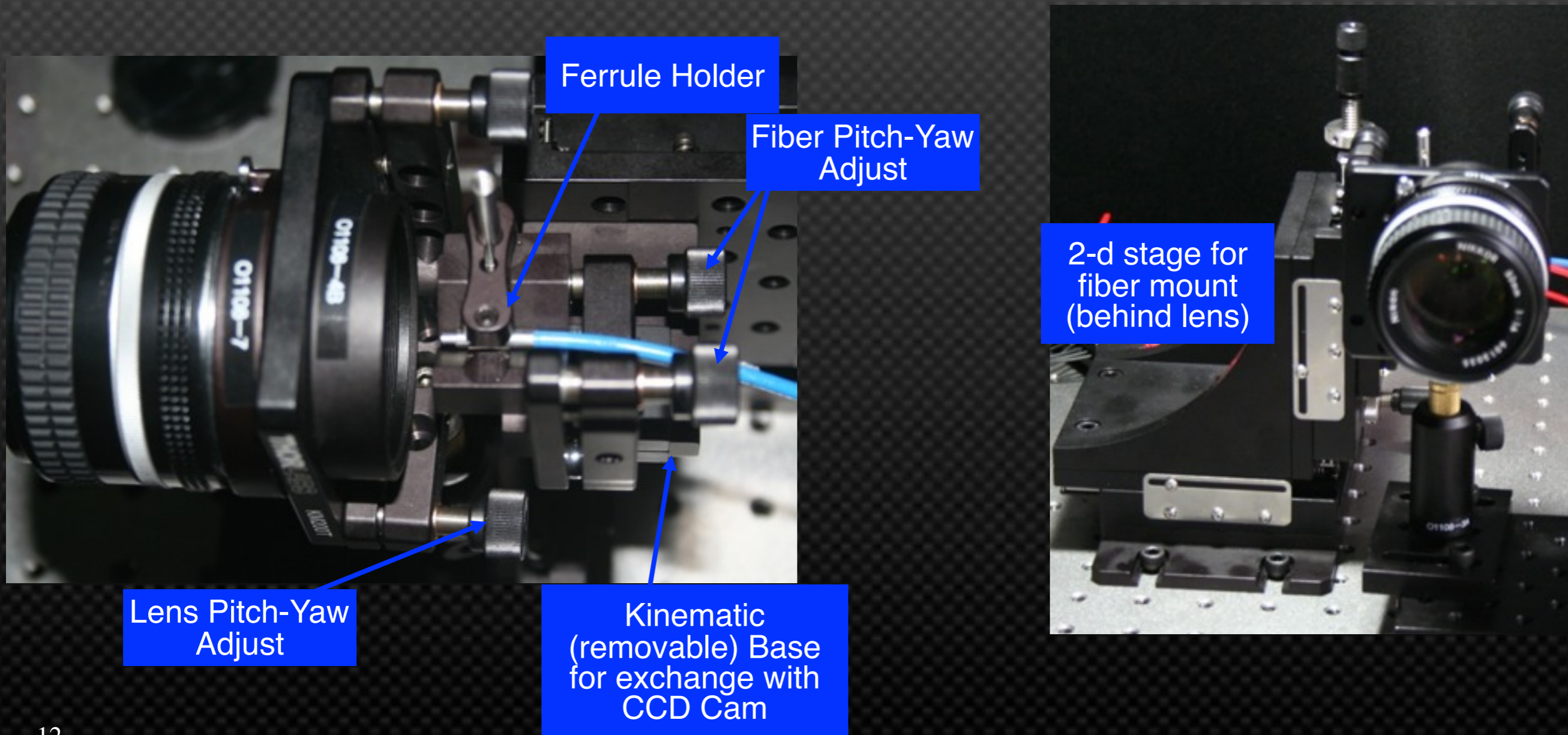
- Measure Output Profile on CCD



# 2011-V3 Setup

- 2011 Hardware

- fibers are difficult to hold, need some special mounts
- alignment of small fibers requires some special mounts and thinking about setup.



# 2011-V3 Setup

- Development of Align Techniques
  - Laser "cross-hairs" + Reflections for lenses
  - Different color laser back-illumination for fiber align



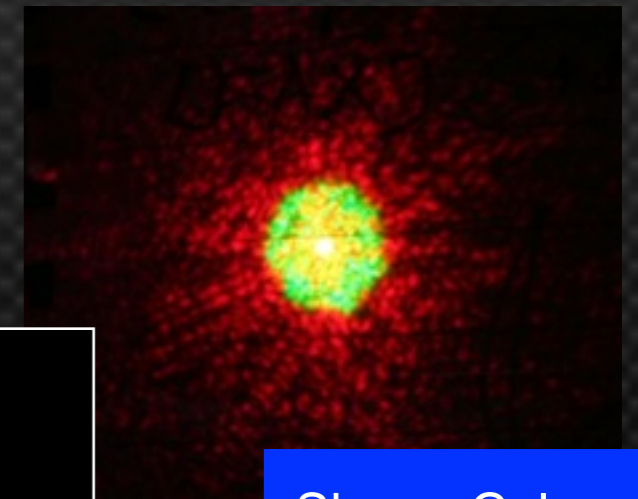
Original  
"Custom" Align-  
o-meter

Reflection off  
Camera Lens  
off-center

Align  $\leq 0.2^\circ - 0.3^\circ$ /lens  
 $\leq 0.05^\circ$  with splitter cube  
... better w/ new screen.



Green Spot  
Aligned On  
Incoming Beam

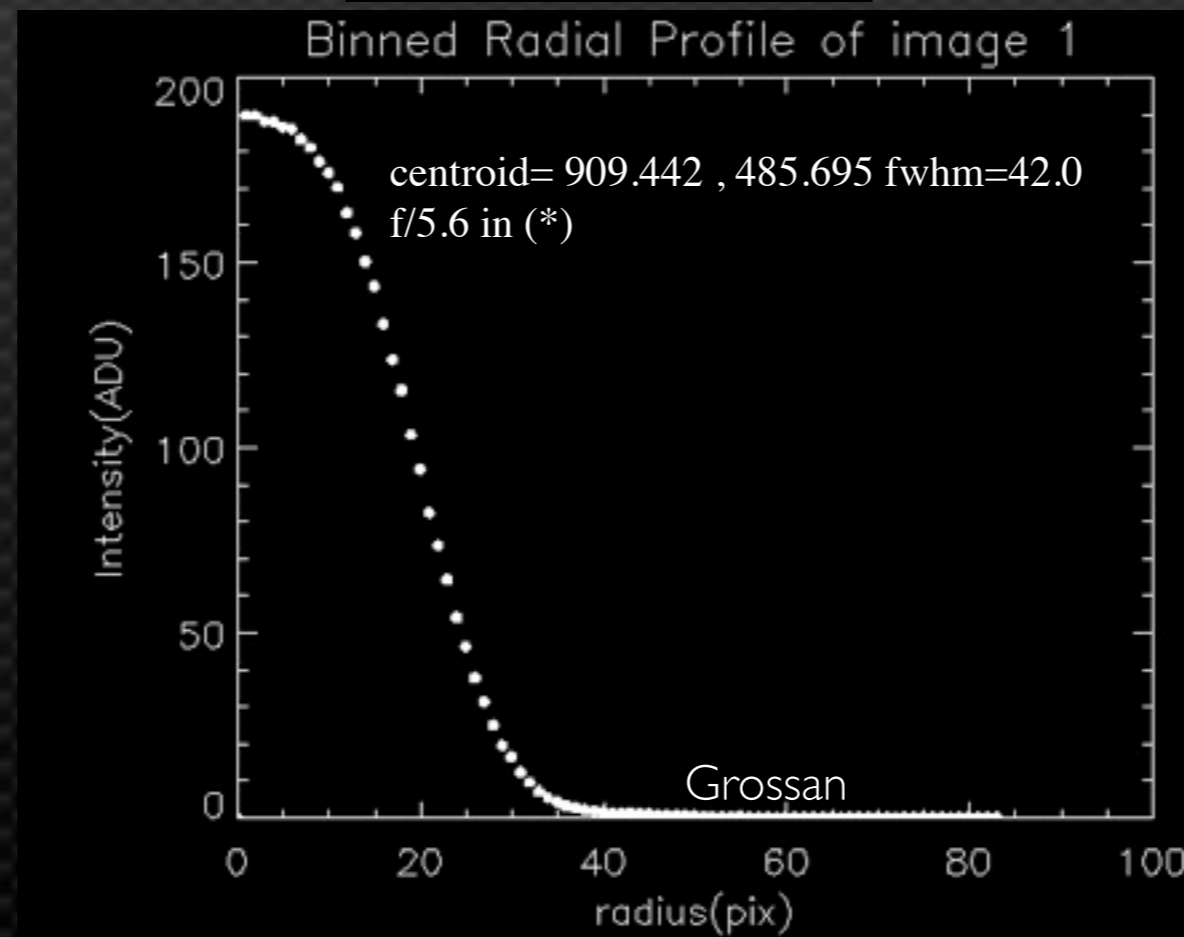
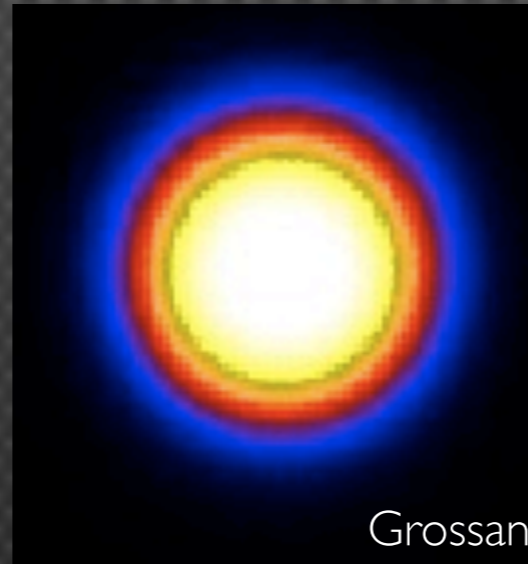


Shows Colors  
Better (spot too  
large though)

Thanks to M. Sholl for helpful suggestions.

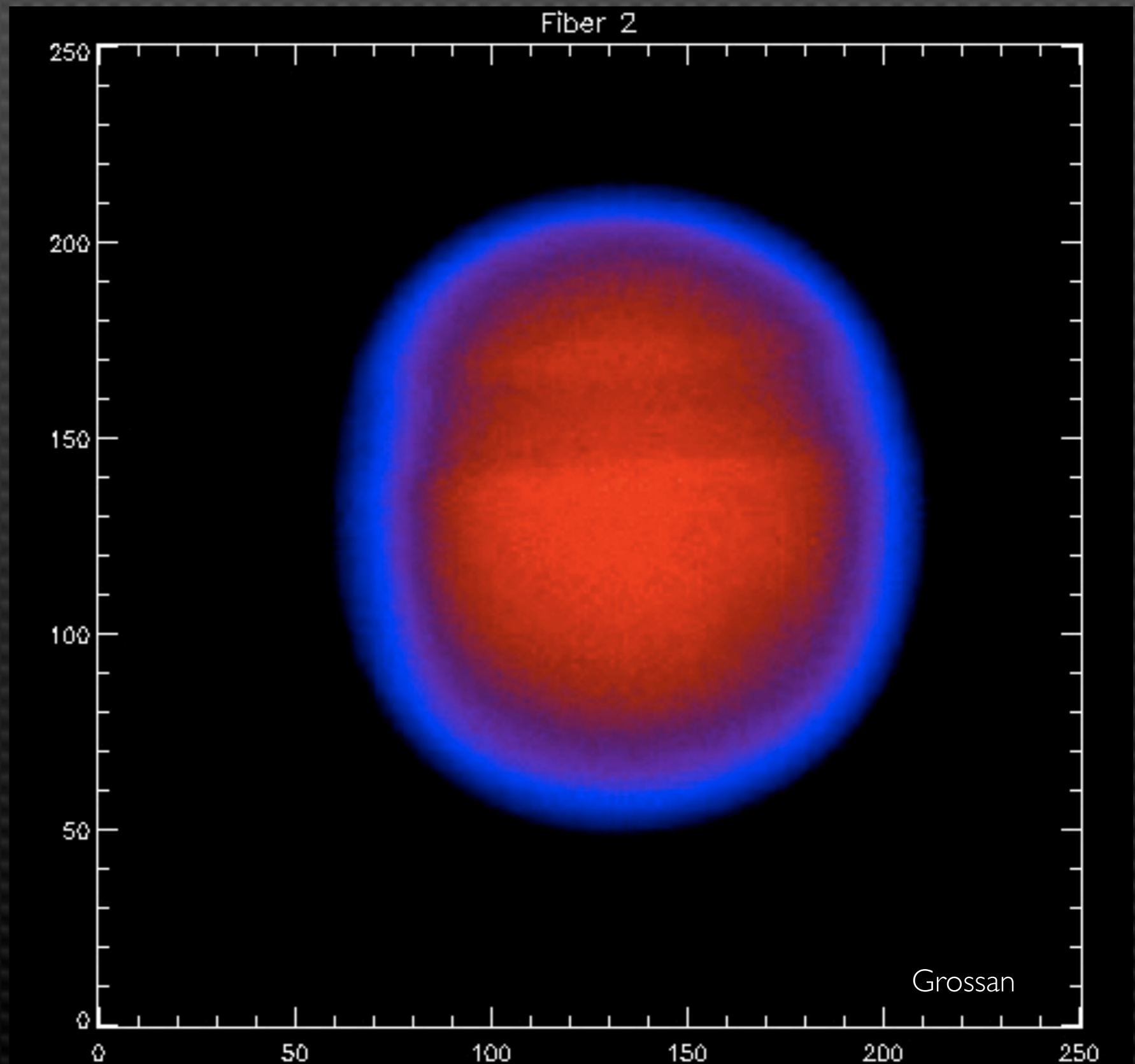
# Results....

- Incoherent Source Profile



# Bad Fibers

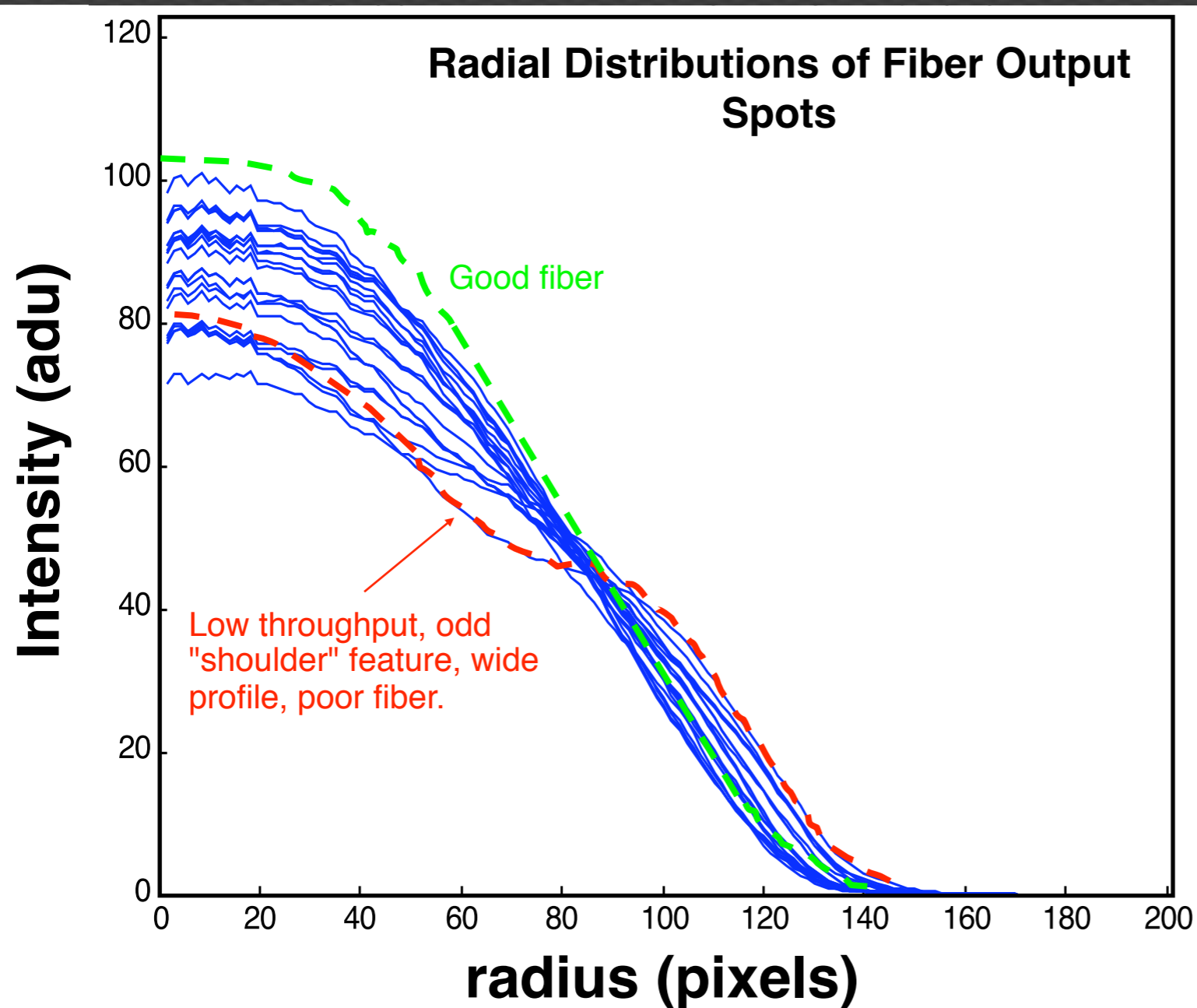
- "Oh no, the alignment must be screwd up?"
- NO! - Fiber 2 has an obviously non-round profile!





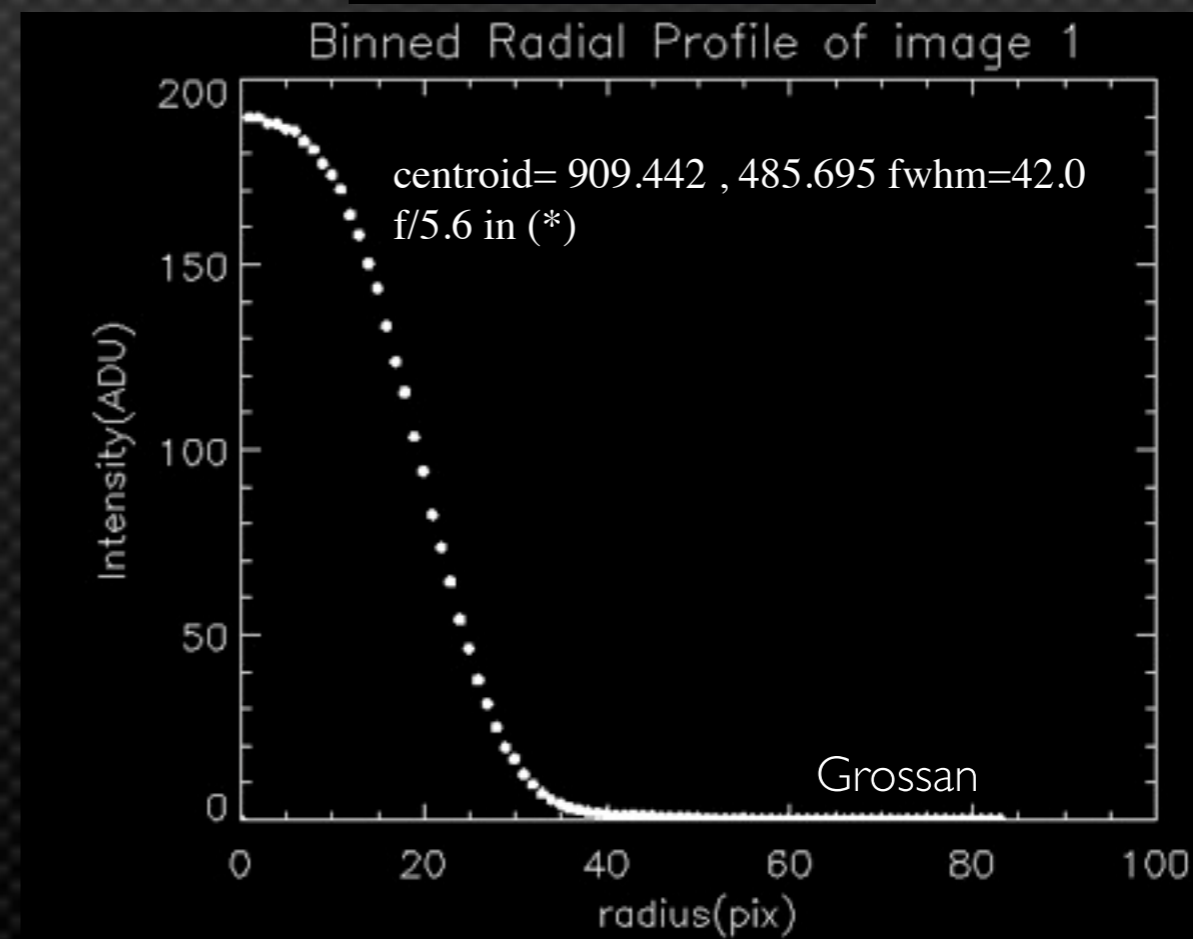
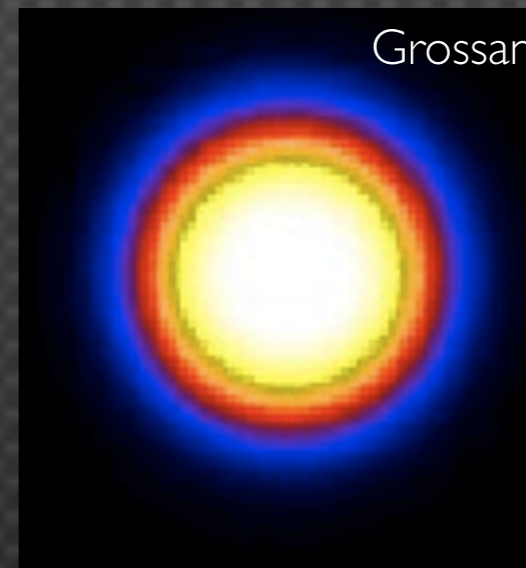
# Comparison

- Tests on BOSS 20-fiber bundle (rejected)
  - Variety of output profiles observed
    - Note: one fiber had zero output



# How do you measure?

- How do you turn this profile into an f/#?
- Please write the program



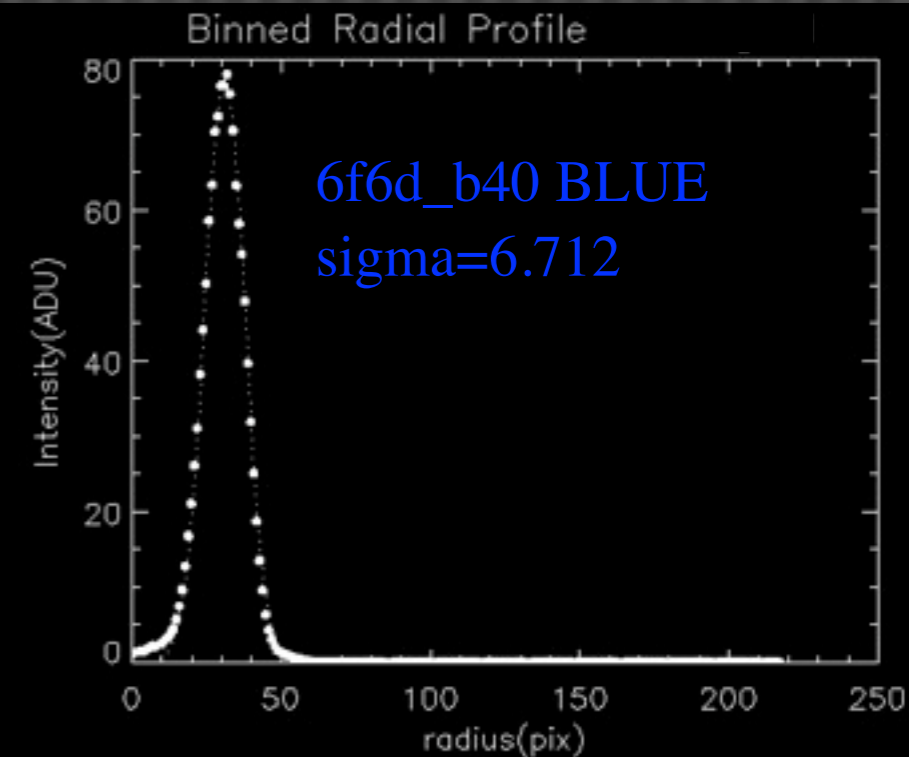
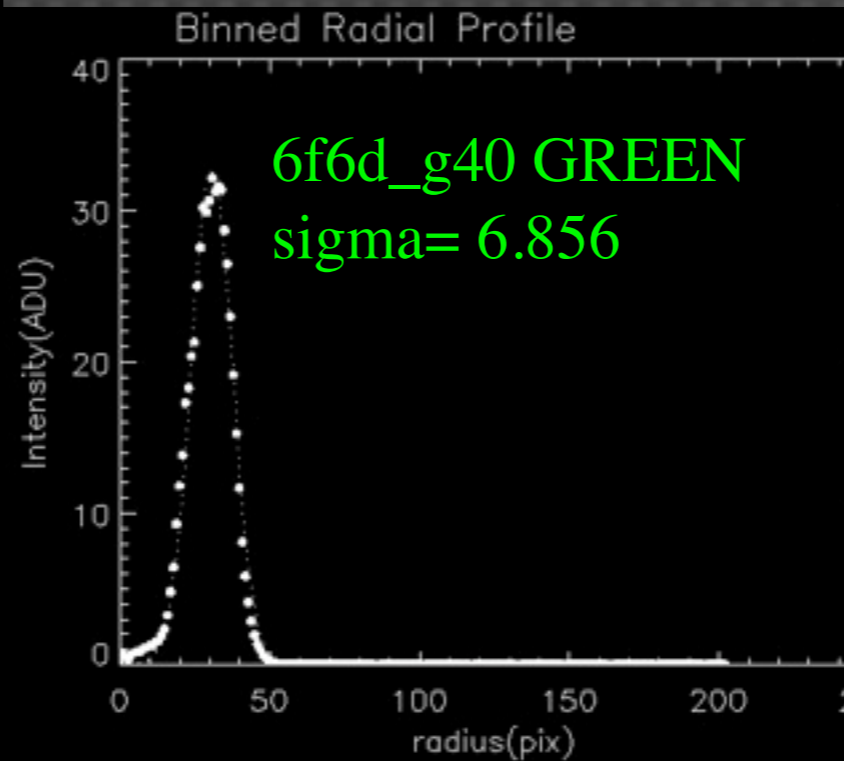
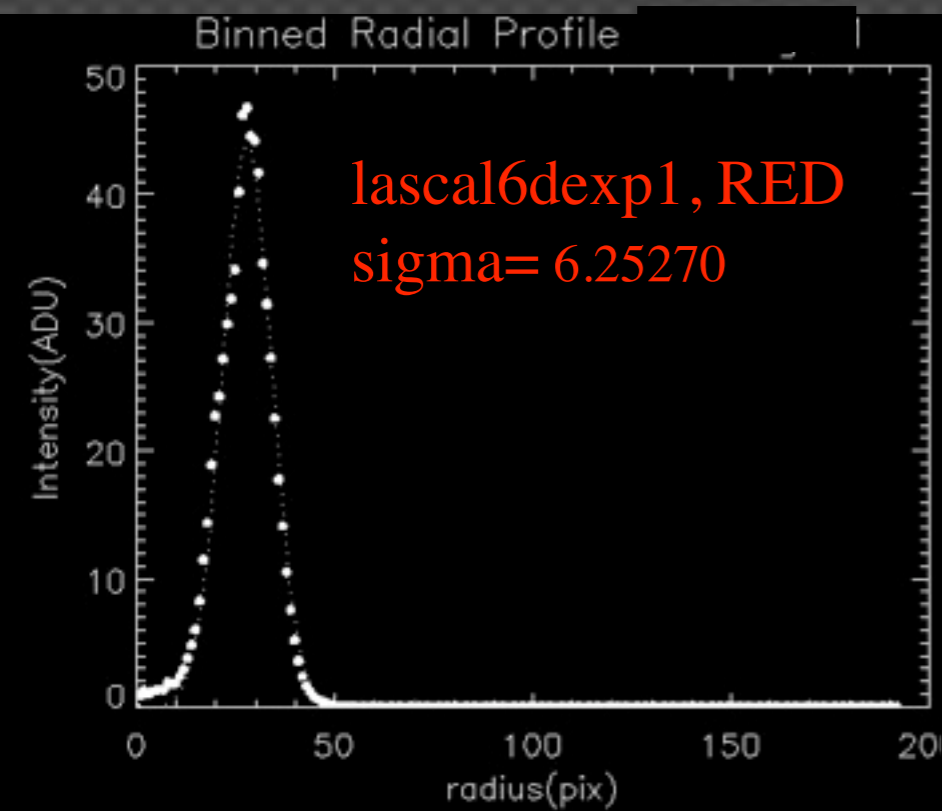
# Color Effects

- Easy to do with diode lasers



# Now Checking Color Effects

- Diode laser test FRD performance for 3 wavelengths.



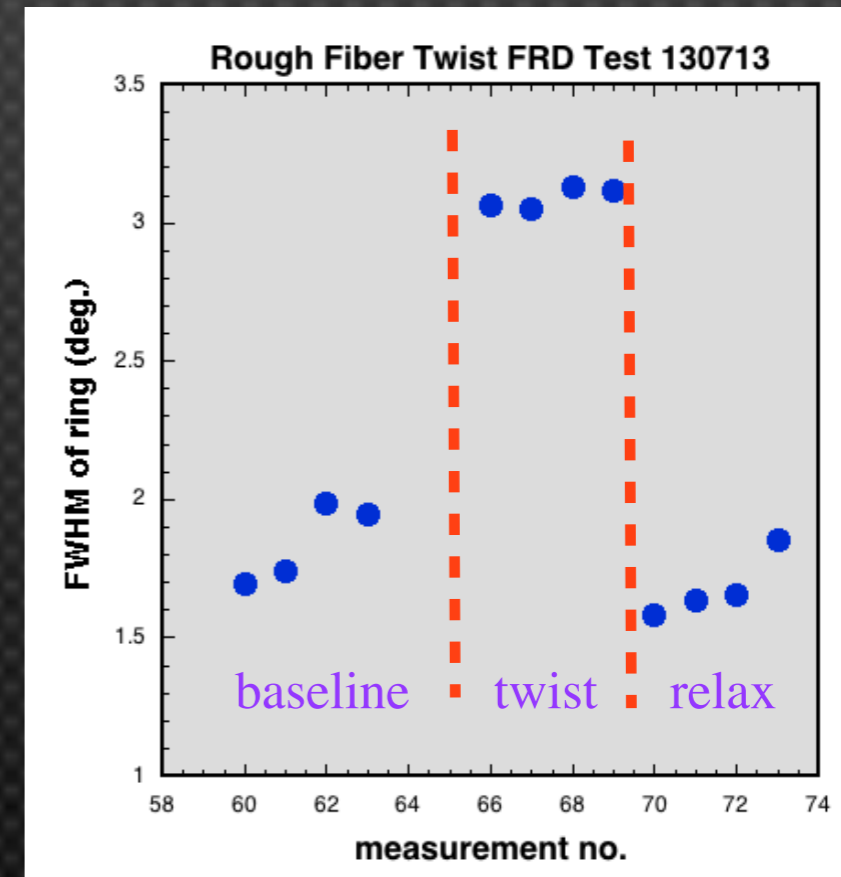
Fiber	Color/ incident angle	sigma fit width
Fib 6	Red/6 deg	6.25
Fib 6	Green/6 deg	6.85
Fib 6	Blue / 6 deg	6.71

**preliminary** -  
 Errors? Repeatability?  
 Varies with fiber?  
 $\lambda$  dependence...?  
 ...working on it now.

Please note "6 deg" = label, only **rough** value of incident angle

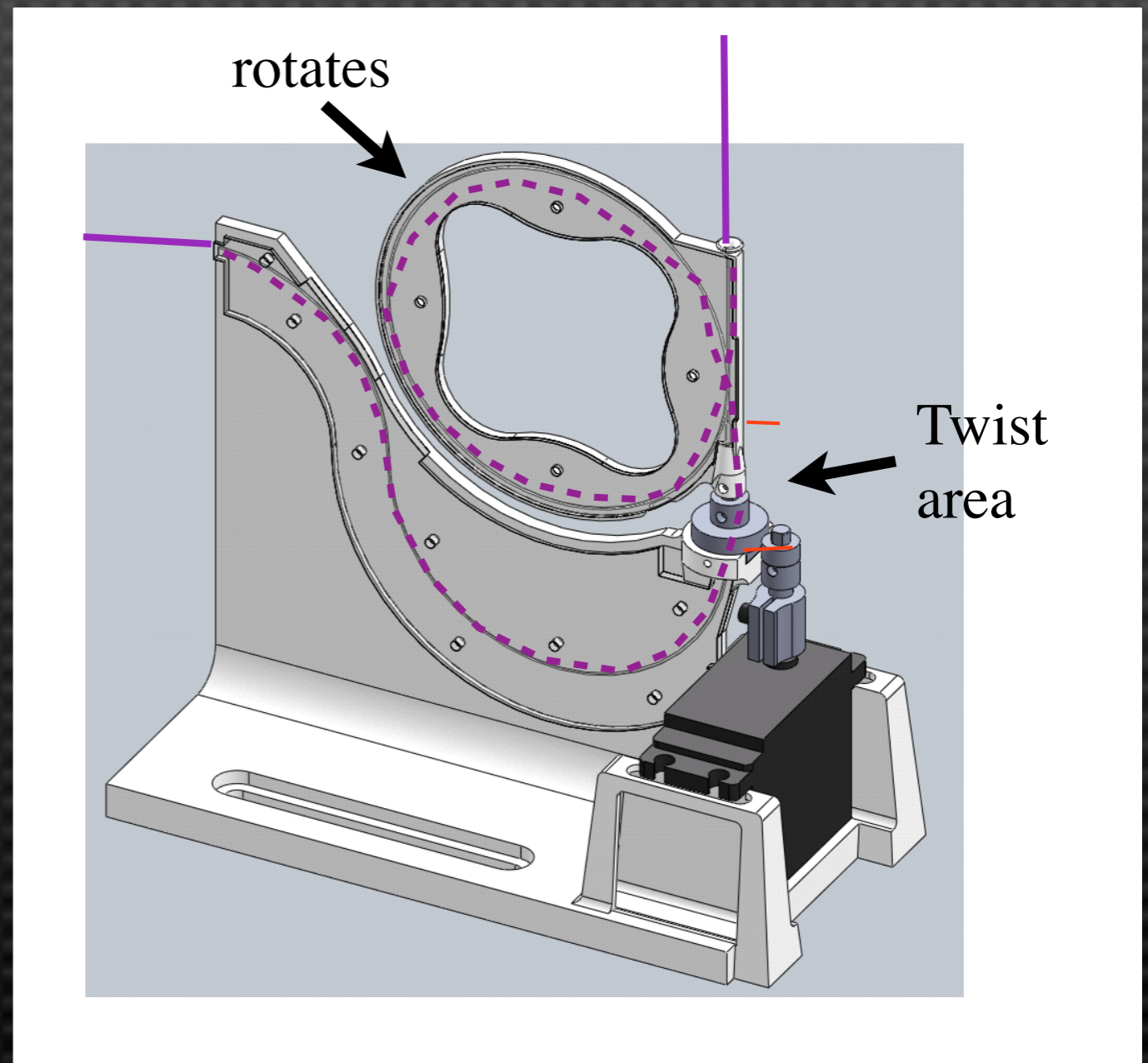
# What about twisting?

- After crude test (pre-twister) of 360 deg. twist in 10 cm, severe FRD results.
- **But is this result of stress at holder, or twist?**



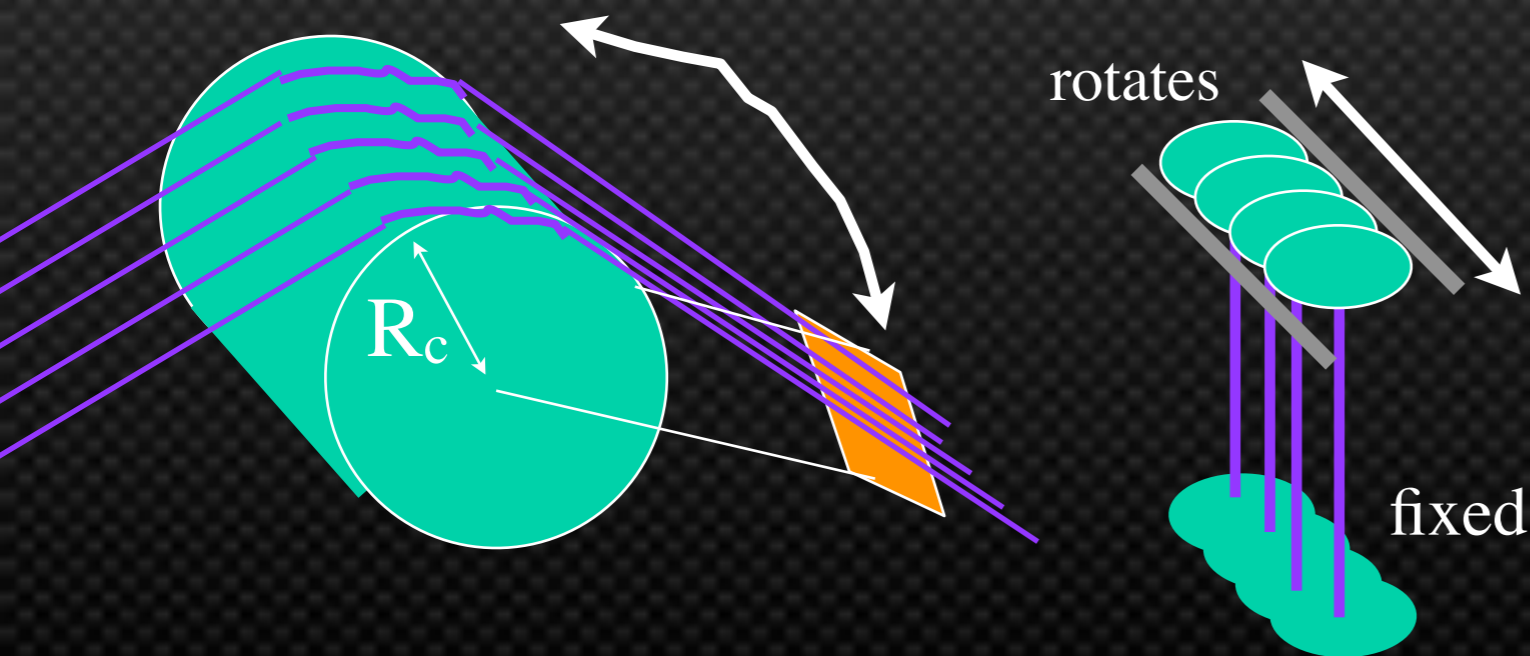
# Twist Testing Device

- Meijer-Grossan twister design
  - Isolates twist so you can learn about it independently.



# Fatigue Testing

- Sholl "Spindle-Board" tests "realistic" bend + twist
  - can mimic theta-theta or r-theta motions
  - easily reproduced in large numbers
  - (thus far guessing on curvature)
- Simple tests separate bend & twist
  - bend over fixed radius of curvature
  - mount fibers in disks, move disks synchronously

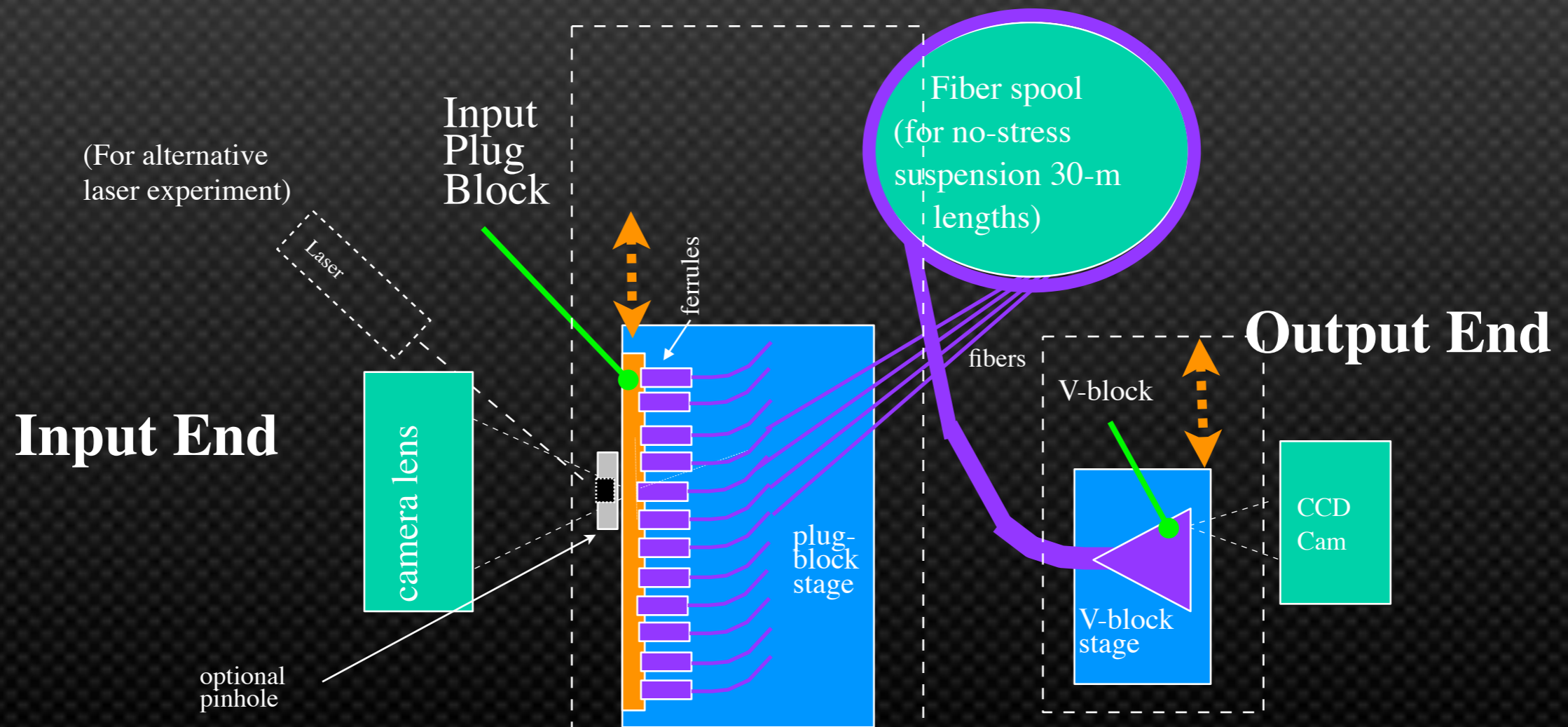


Cheap, scalable, programmable micromotors & controllers by Sholl.

- Test large samples using IEU people/facilities

# Next Step: Automate for 5000 tests

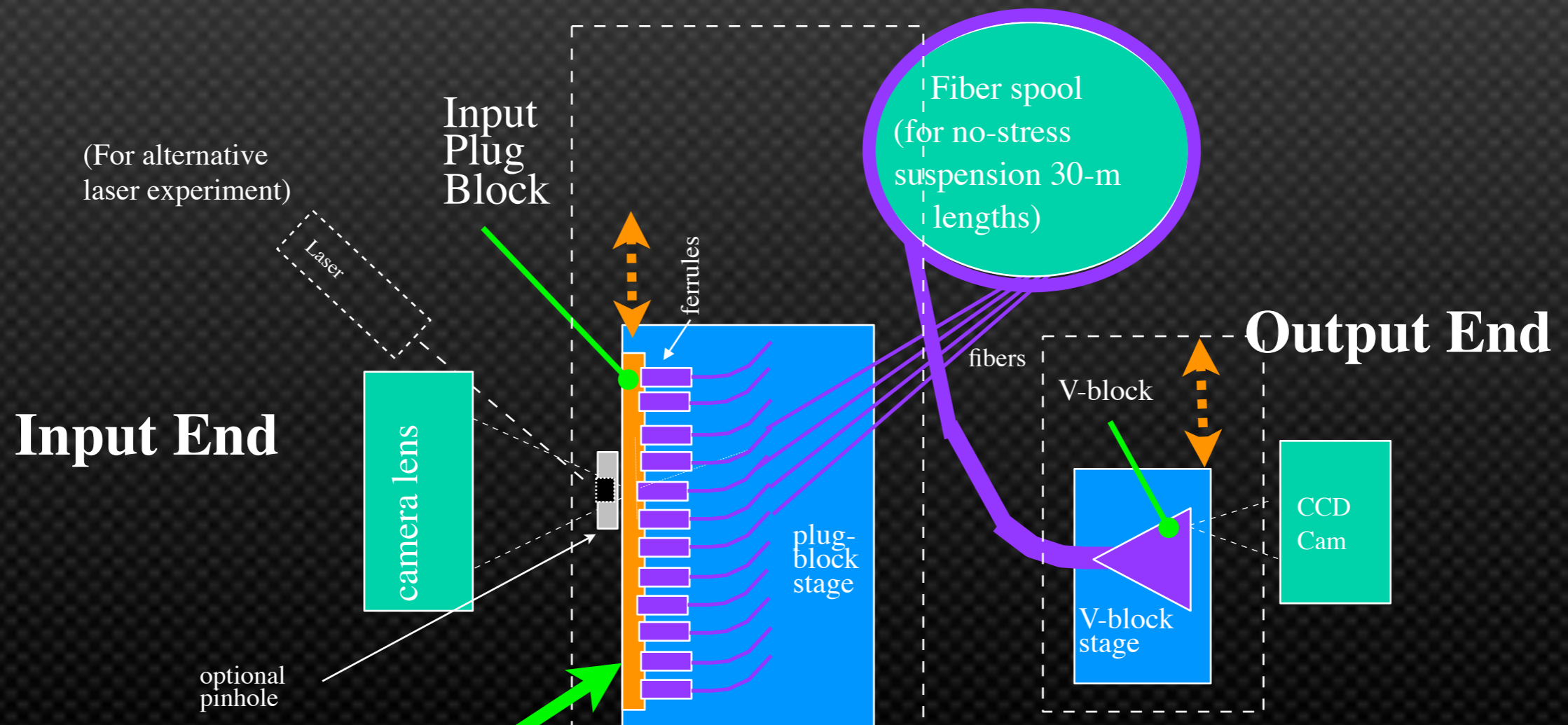
- replace plug block(input end), V-block (output end)holder with 1-d computer-controlled stages
  - align first and last fiber on block
  - scan over each (some?) fiber(s) to find align point @ max intensity.
  - 2011: **Test Closed-Loop Feedback System**





# Next Step: Automate for 5000 tests

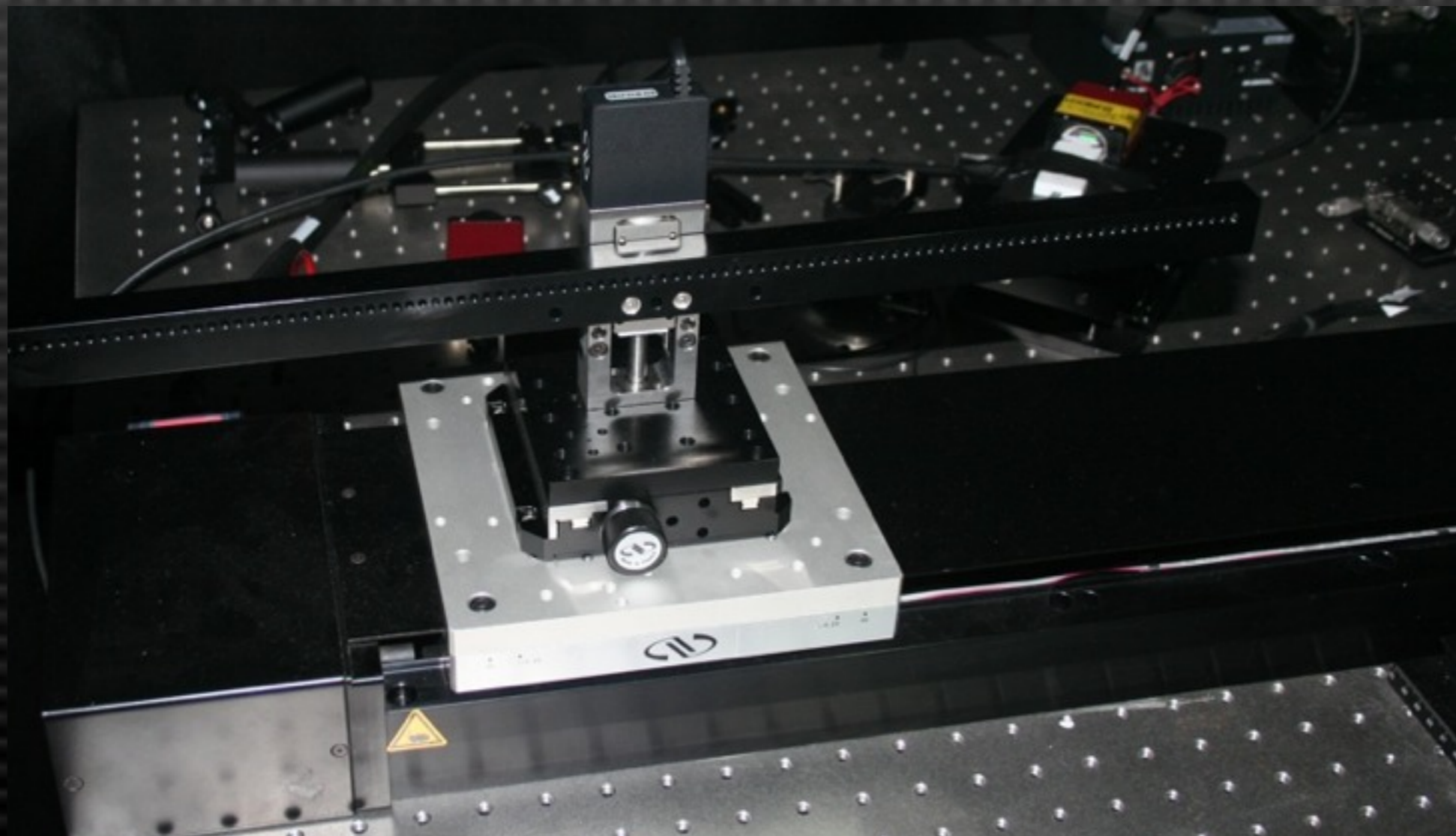
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This plug block is very long (because many fibers). Need long-travel stage!

# Computer Controlled Stages

- **Challenge is alignment**; cannot adjust thumbscrew for every fiber
- **Currently Validating**  $\mu\text{m}$  stage performance: how bad hysteresis during max. align scans? Accumulation of errors?



# QUESTIONS

- Spectra are recorded as “strips” of light on CCD detectors.  
Let’s say we measured FRD that made the width of a spot of a fiber 2X as wide as that for a perfect fiber.
- What would be the effect on signal-to-noise ratio of this FRD?
- Draw the spectrum of the [OII]3727Å, 3729Å infinitesimally narrow lines, for a spectrograph with 0.3Å FWHM resolution, 0.3 Å pixels, fed by a fiber with a FWHM spot size of 1 pixel (Draw reduced spectrum and draw “strip” on CCD).
- Again, this time with an optical fiber with 5 pixel spot size. Does this fiber affect the spectroscopy? If so, how?

# SUMMARY

- You can use optical fibers, positioned in the places where light from galaxies falls on a telescope focal plane, to take multiple spectra in a single exposure. This allows you to do millions of objects spectroscopic surveys
- Fibers have FRD, a kind of imperfection which causes light coming through the fiber to spread out over a large area.
- Shining coherent light into a fiber makes a ring - the width of that ring tells you how much the fiber is spreading out the light.
- If you build a fiber spectrograph, make sure your fiber performance is good enough (FRD small enough) to get the answer you need!

# END

- THANKS!