

Space environment: effect on detectors examples with Integral and Astro-H

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Plan

1. Description of the space environment.
2. Effect of radiation on electronic devices.
3. Calibration project for Astro-H.

The space environment

Space constraints

➤ Launch

- Chocks, vibrations

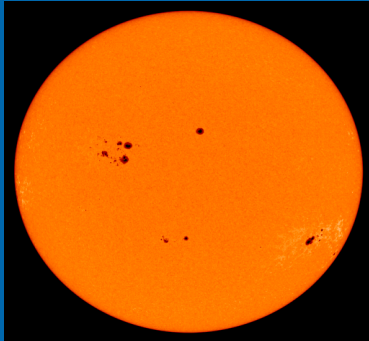
➤ Space environment

- Thermal Variations
- Vacuum
- Radiations
- Contamination
- Micrometeorites, space debris

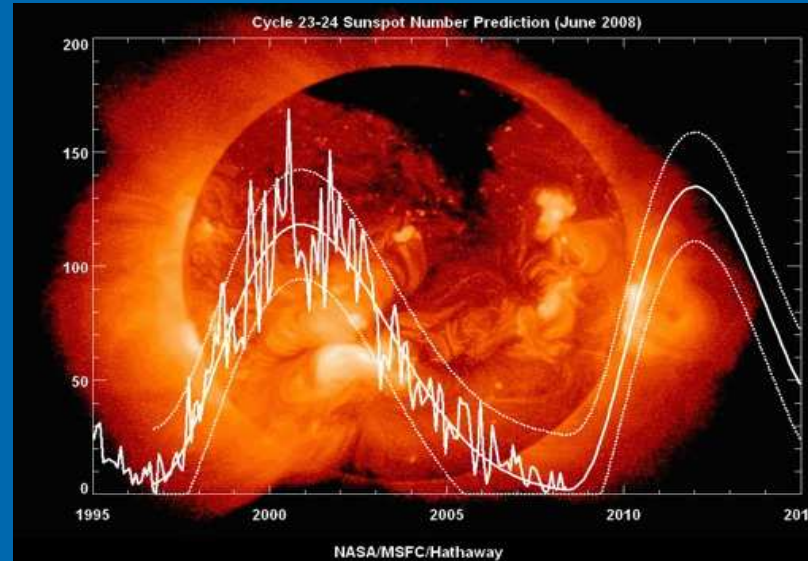
The radiative environment

Solar activity

Sunspot
(Galilée, 1610)

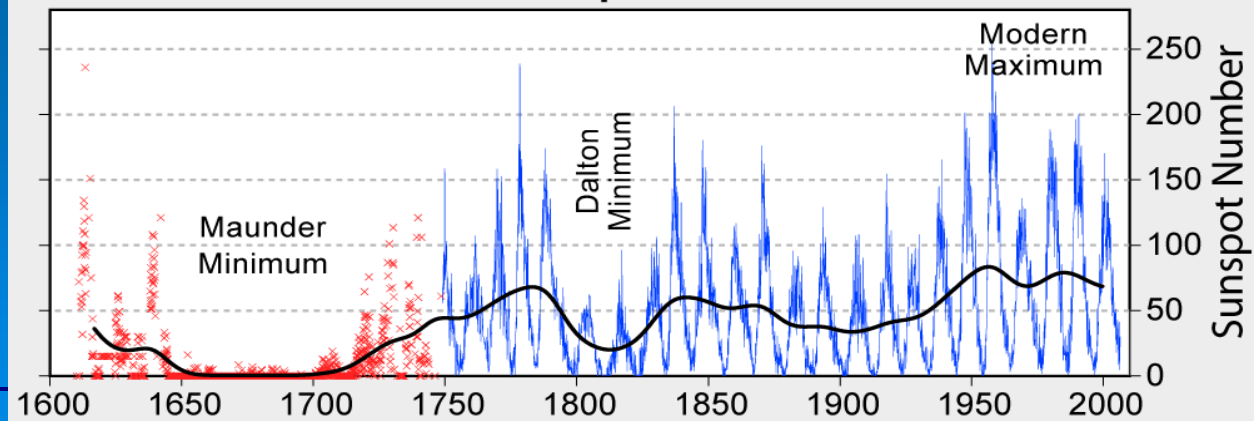


flares



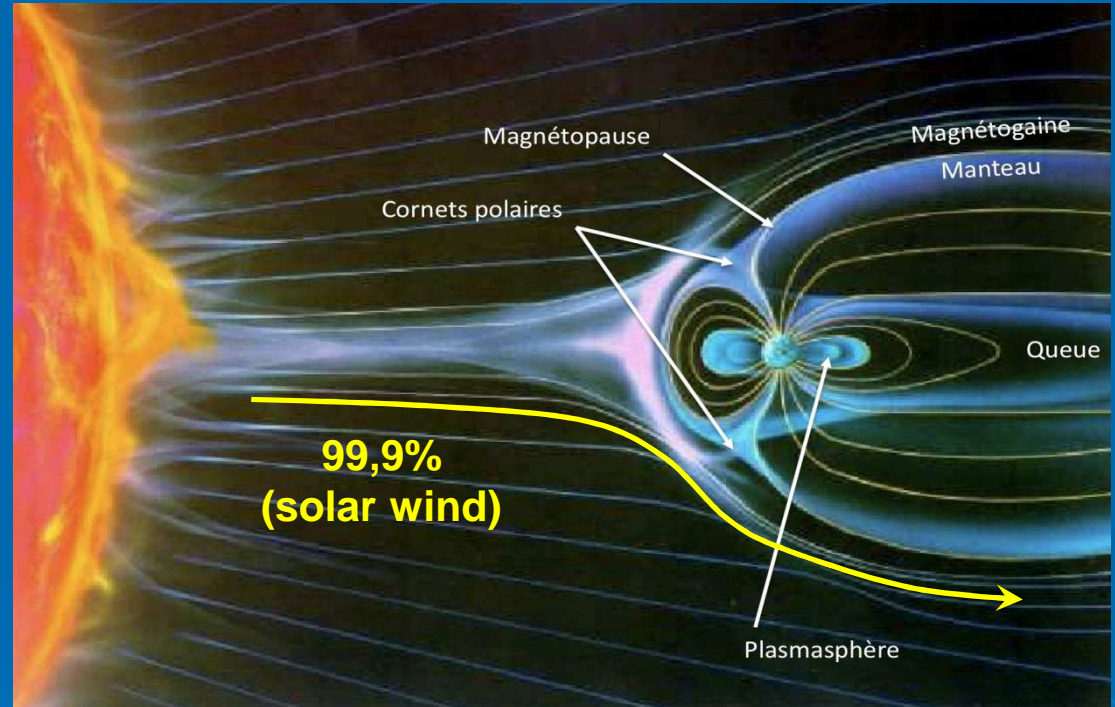
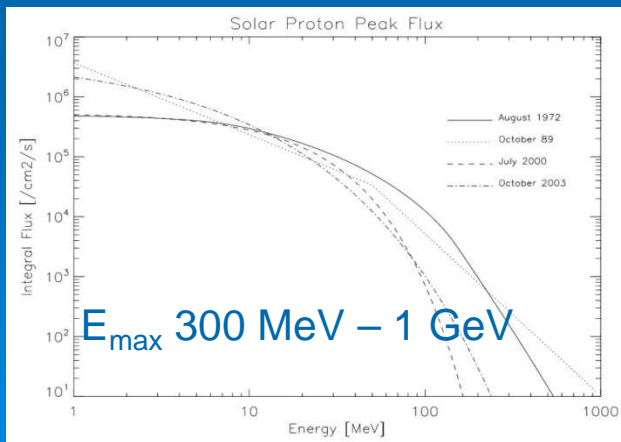
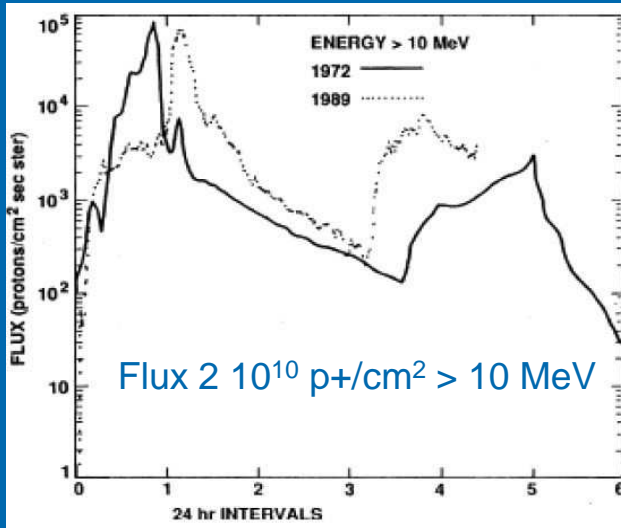
11 ± 2 years cycle
Cycles are not similar !

400 Years of Sunspot Observations

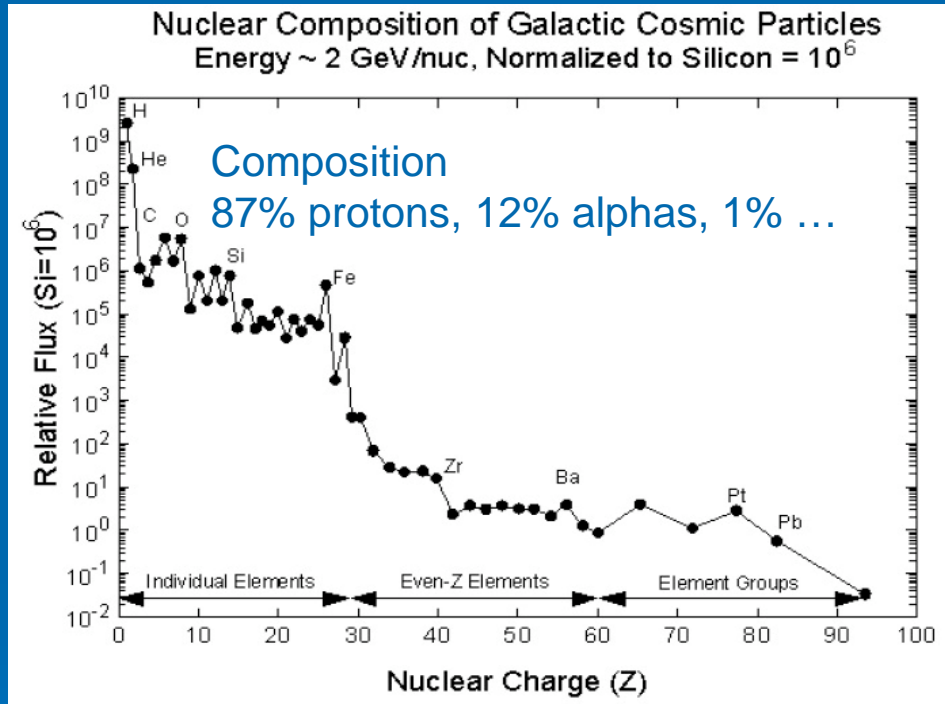


Solar particles

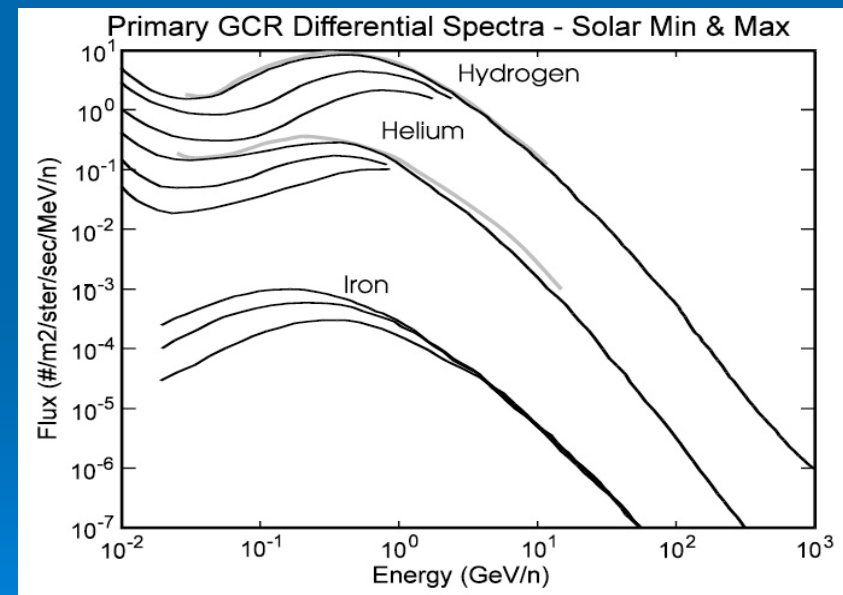
Events occurring on hours to days timescale



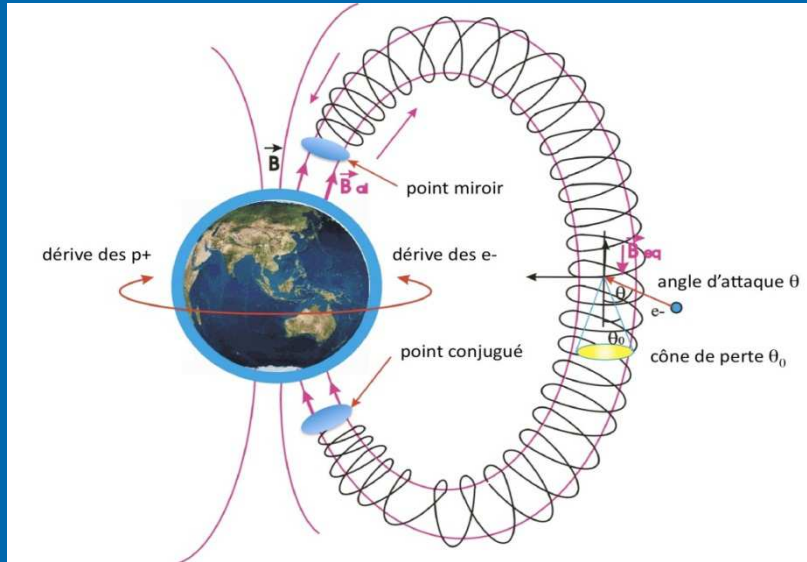
Cosmic Rays



Solar modulation

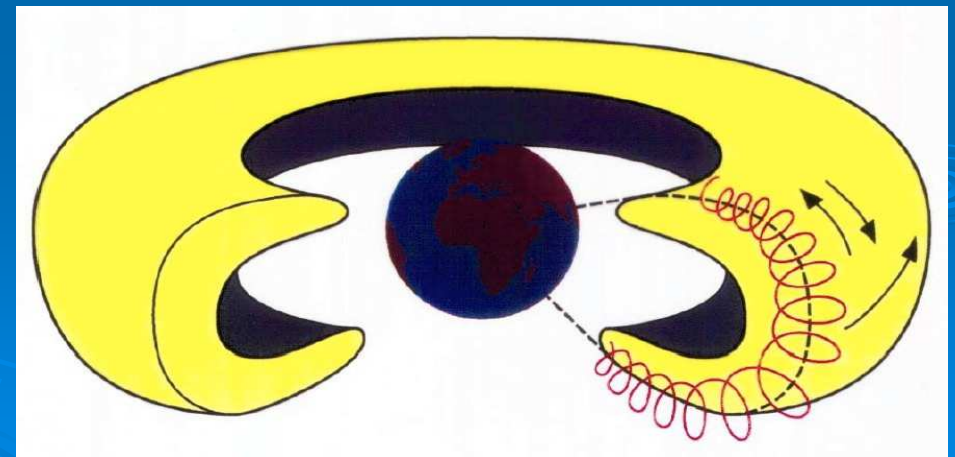
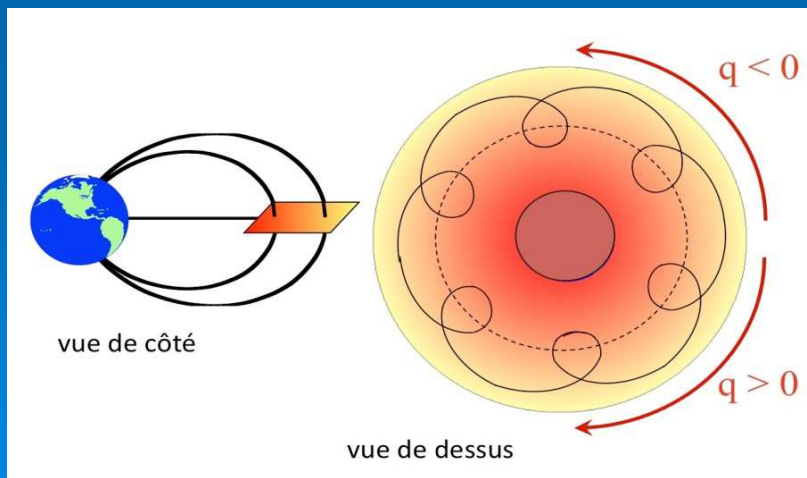


Trapped particles



Trapping mechanisms:

- rotation (Lorentz force)
- Magnetic mirror.
- East-West drift, due to magnetic field gradient.

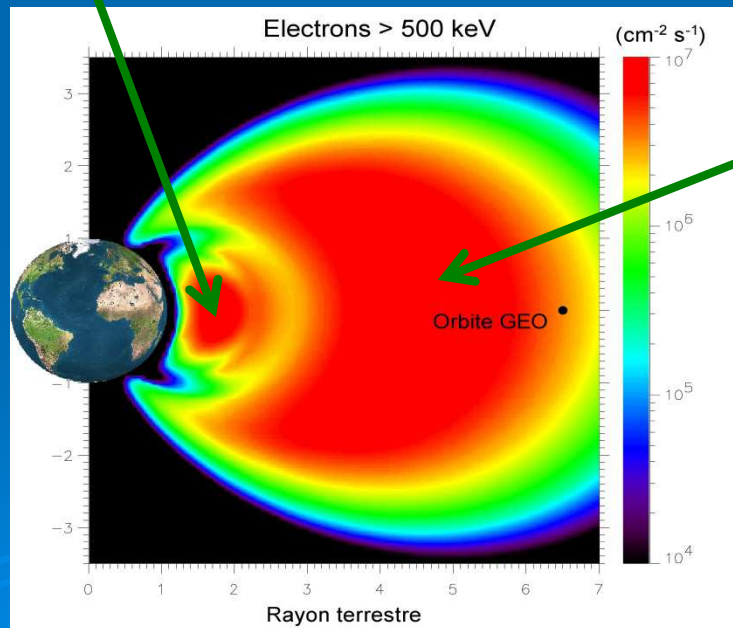
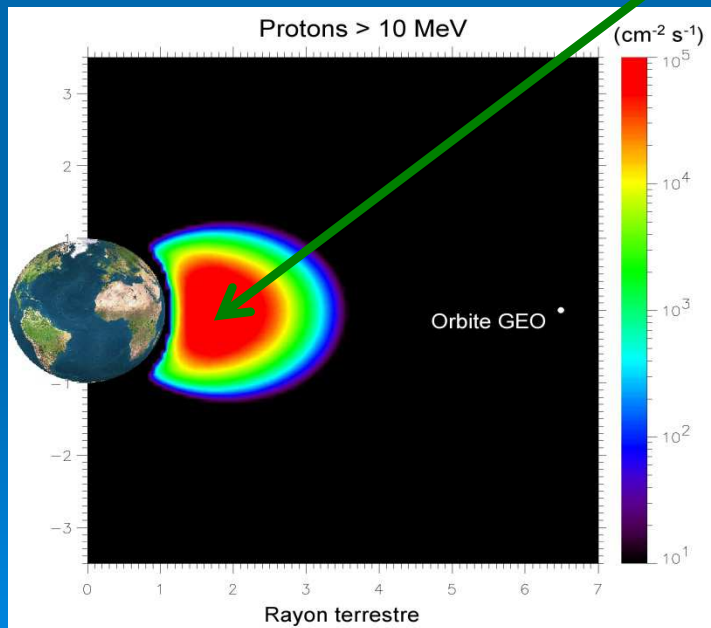
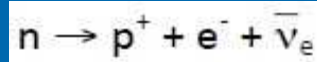


Van Allen belts

Two belts:

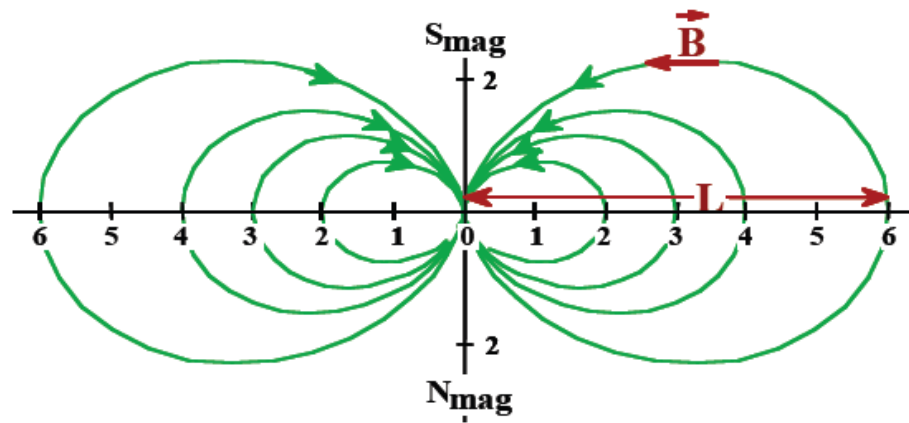
- Internal (stable)
- External (dynamic)

Protons and electrons coming from cosmic and solar particles interaction into the atmosphere



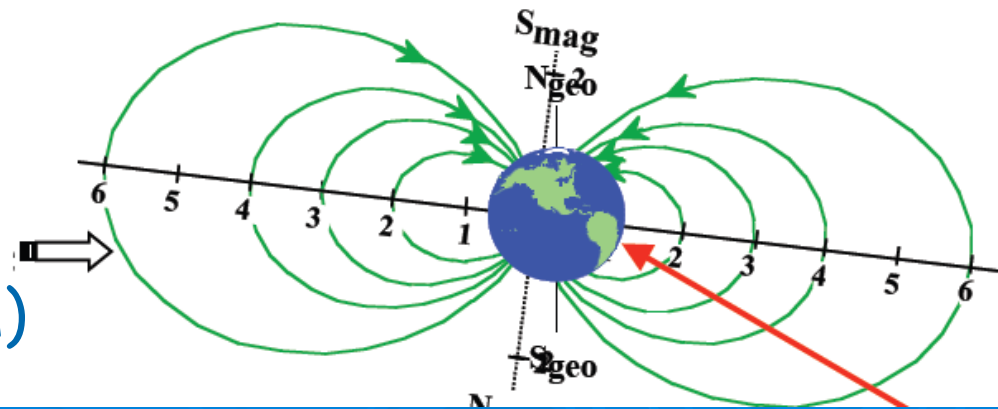
Solar particles injected from the magnetosphere tail.

South Atlantic Anomaly (1/2)



Dipole field lines

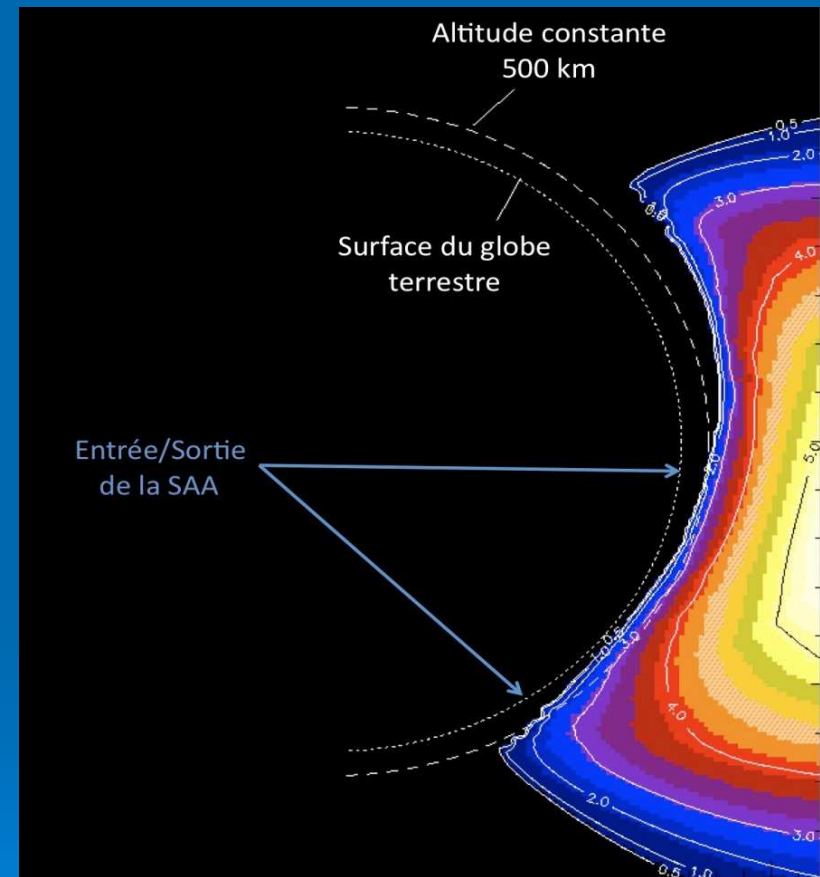
- Earth magnetic field:
- inclined ($\sim 10^\circ$)
 - Center offset (~ 500 km)



SAA

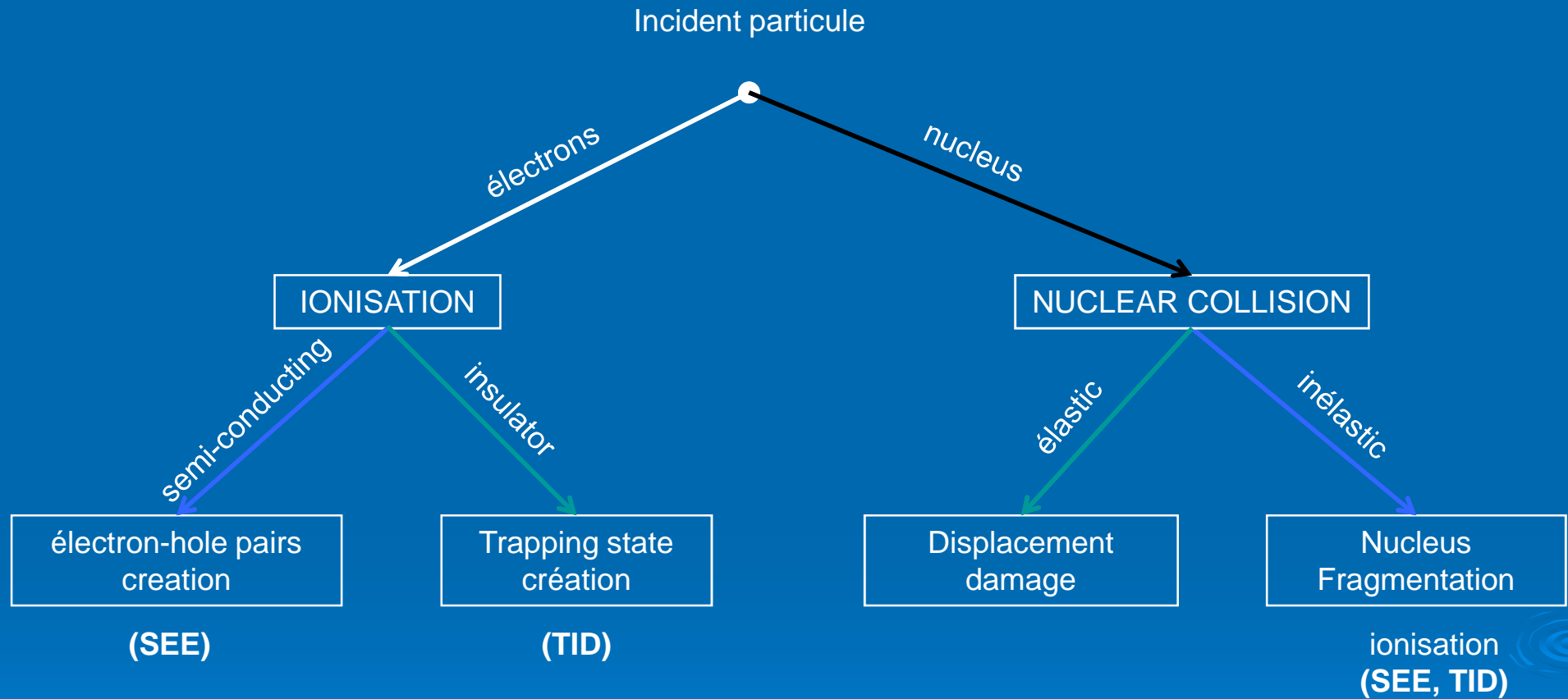
South Atlantic Anomaly (2/2)

- The **South Atlantic Anomaly (SAA)** is an area where the Earth's inner Van Allen radiation belt comes closest to the Earth's.
- This leads to an increased flux of energetic particles in this region.
- The effect is caused by the non-concentricity of the Earth and its magnetic dipole.



Effect of radiations on electronic devices

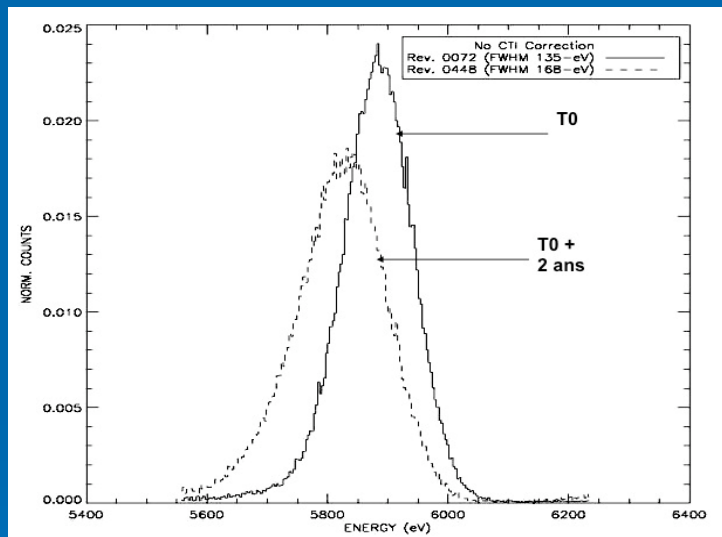
Effects on electronic devices



- TID : "Total Ionising Dose" is a long cumulative effect
- SEE : "Single Event Effect" is a transient single effect

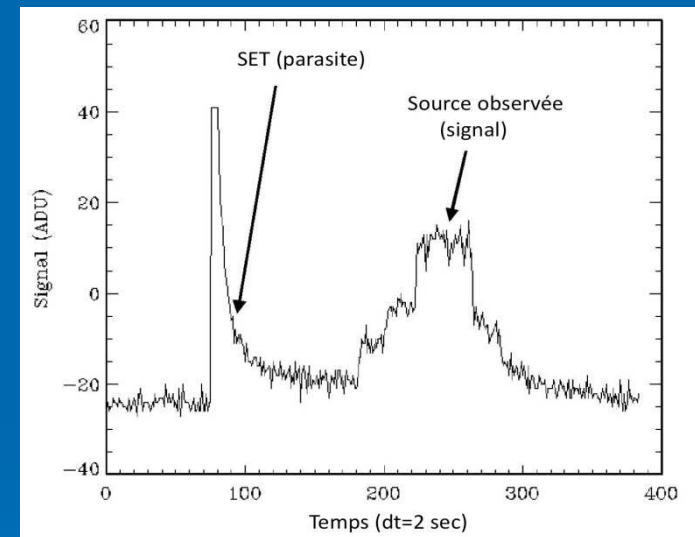
Examples

Dose →



Spectral resolution dégradation (XMM)

SEE →



Transient signal (ISOCAM)

Effect of SEE on electronic devices

- Effect of the ionizing particle depends on the energy deposit in the component (LET):

$$LET = \frac{1}{\rho} \frac{dE}{dx}$$

- **Latch-up (SEL)** : parasitic current induced by the particle passage which could produce a short-circuit and potentially destroy the component.
- **Single event upset (SEU)**: bit flip in a register.
- Radhard Technics to immunize the electronics against SEE.

Component tolerance to SEE (1/3)

Measures of the circuit properties

SEL and SEU (*single event latch-up / upset*) cross sections with respect to LET (*linear energy transfer*):

$$\sigma = \frac{N_{LET}}{N_{inc}}$$

N_{inc} = number of incident ions which will deposit a given LET.



M/Q = 3,33

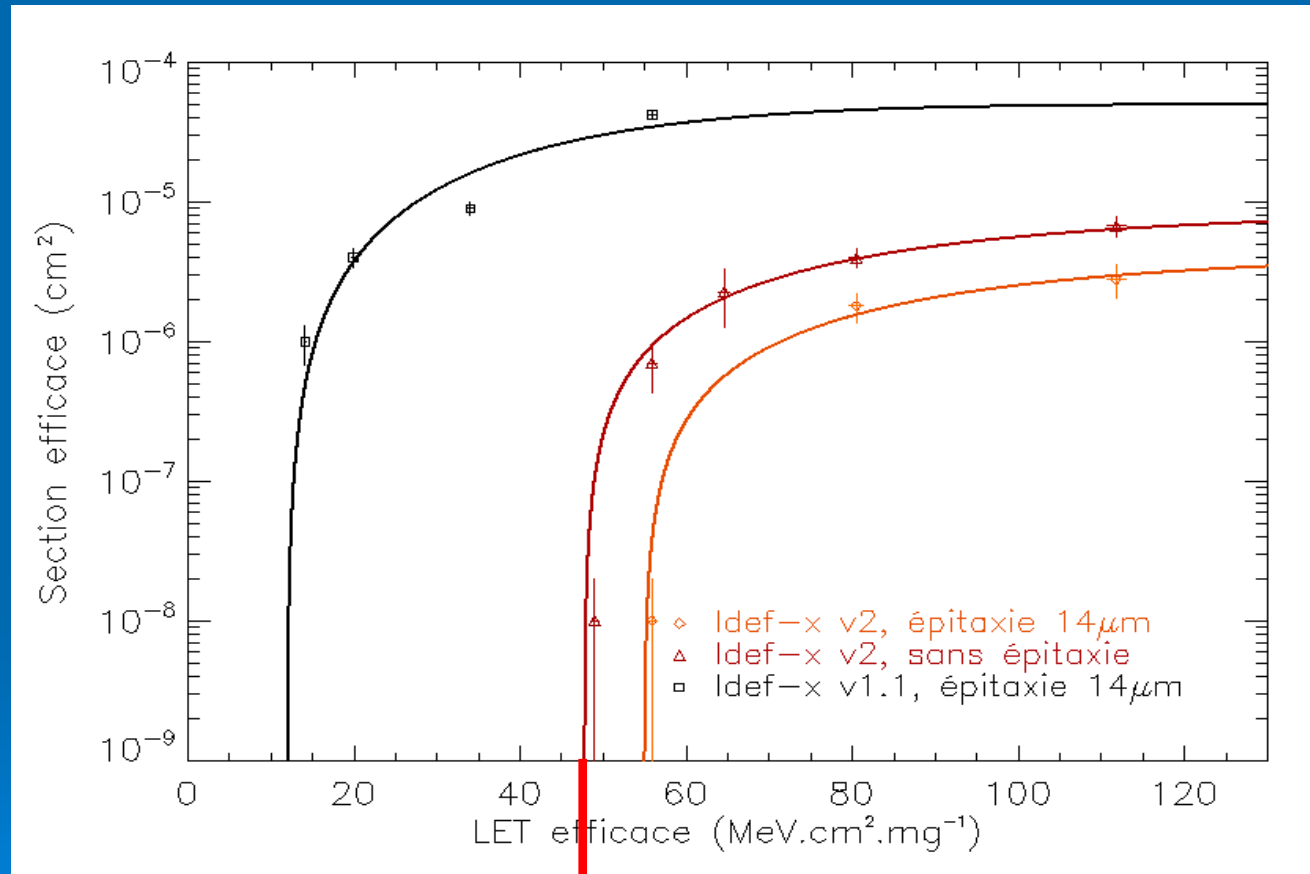
| Ion | Energy [MeV] | LET [MeV/mg/cm2] | Range [microns] |
|------------------------|--------------|------------------|-----------------|
| $^{13}\text{C}^{+4}$ | 131 | 1.2 | 266 |
| $^{22}\text{Ne}^{+7}$ | 235 | 3.6 | 199 |
| $^{40}\text{Ar}^{+12}$ | 372 | 9.95 | 119 |
| $^{58}\text{Ni}^{+18}$ | 567 | 21.3 | 98 |
| $^{83}\text{Kr}^{+25}$ | 756 | 31.0 | 92 |

M/Q = 5

| Ion | Energy [MeV] | LET [MeV/mg/cm2] | Range [microns] |
|-------------------------|--------------|------------------|-----------------|
| $^{15}\text{N}^{+3}$ | 62 | 3.3 | 64 |
| $^{20}\text{Ne}^{+4}$ | 78 | 6.2 | 45 |
| $^{40}\text{Ar}^{+8}$ | 150 | 15.9 | 42 |
| $^{84}\text{Kr}^{+17}$ | 316 | 40.1 | 43 |
| $^{132}\text{Xe}^{+26}$ | 459 | 67.7 | 43 |

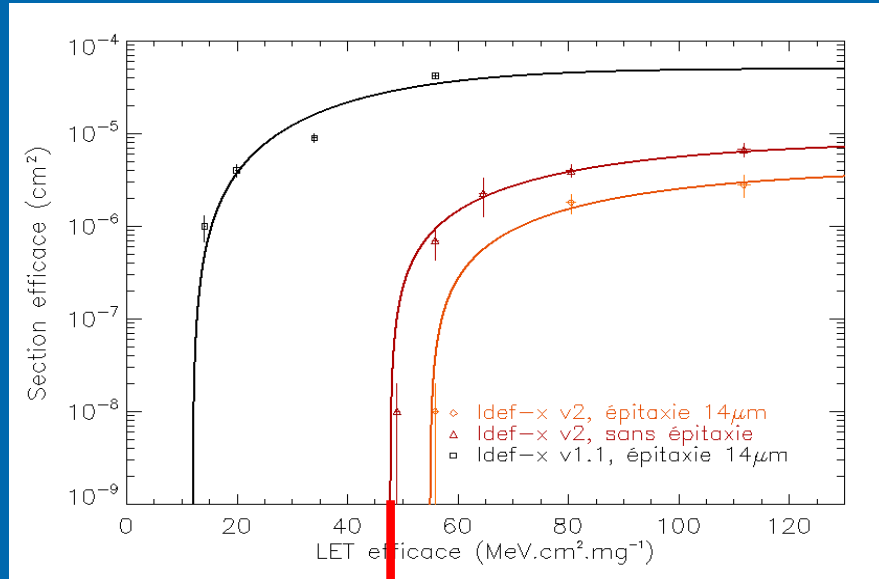
Component tolerance to SEE (2/3)

Measures of the circuit properties



LET threshold

Component tolerance to SEE (3/3)



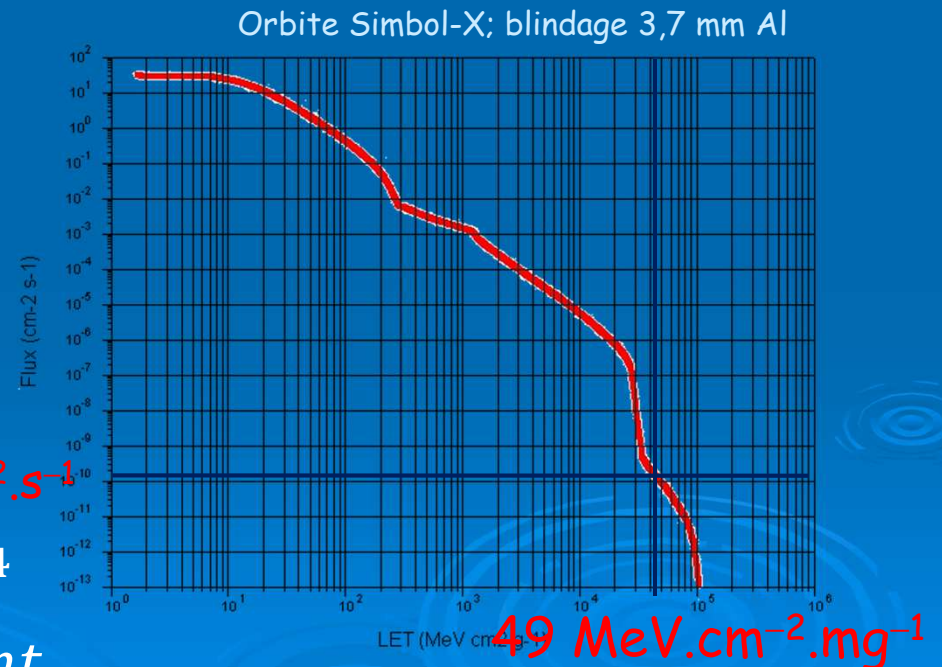
49 MeV.cm⁻².mg⁻¹

10⁻¹⁰.cm⁻².s⁻¹

$prob_{SEE} = 10^{-10} A_{comp} t_{orb} = 4 \cdot 10^{-4}$
 for 3 years and 2 x 2 mm component

Failure probability in orbit


- Convolution of the cross section (cm²) by the LET spectrum (cm⁻².s⁻¹) gives the foreseen SEE rate (s⁻¹) :



49 MeV.cm⁻².mg⁻¹

Calibration project for Astro-H

Context: ESA contribution to Astro-H

| | | |
|-----------------------------------------------------------------------------------|----------------------------------------|-----------------------------------------------------------------------------------------------|
|  | ESA CONTRIBUTION TO ASTRO-H | Doc. no. : SRON-ASTH-PL-2010-004 Issue : 2.2 Date : 26 September 2010 Page : 2 of 36 |
|-----------------------------------------------------------------------------------|----------------------------------------|-----------------------------------------------------------------------------------------------|

ESA contribution to the Japanese X-ray mission ASTRO-H

Technical lead:

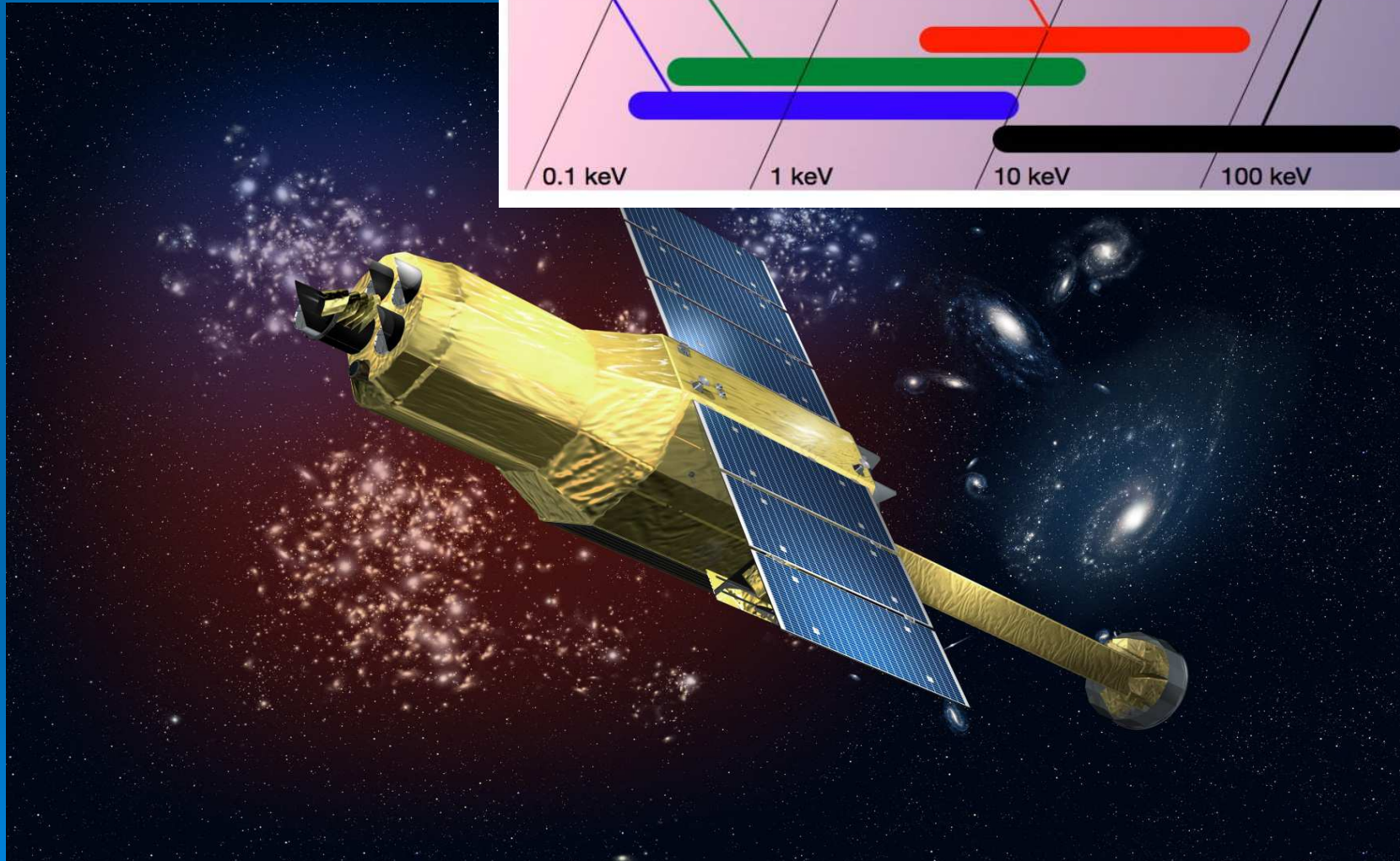
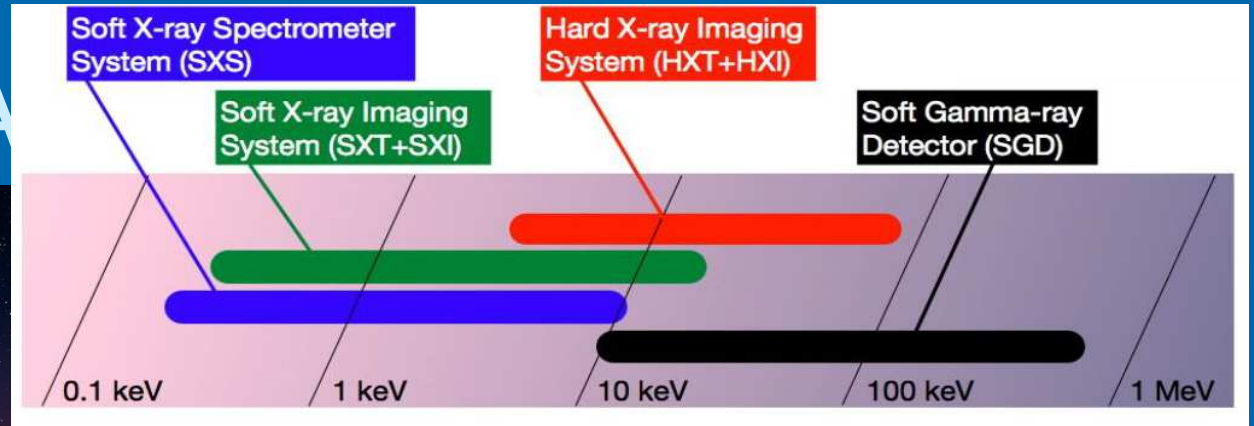
| | |
|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|--------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| Dr. J.W. den Herder (proposal leader) SRON, the Netherlands Institute for Space Research Sorbonnelaan 2 3854 CA Utrecht The Netherlands Tel: +31 88 777 5894 e-mail: J.den.Herder@sron.nl | Dr. P Laurent CEA/DSM/IRFU/SaP 91191 Gif-sur-Yvette Cedex France Laboratoire APC 10, rue Alice Domont et Léonie Duquet 75205 Paris Cedex 13 France Tel: +33 1 69 08 80 66 e-mail: Philippe.Laurent@cea.fr |
|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|--------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|

Context

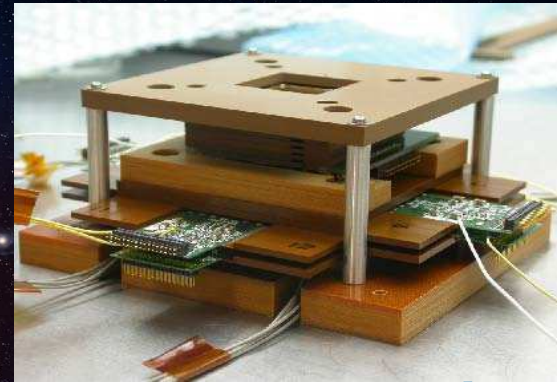
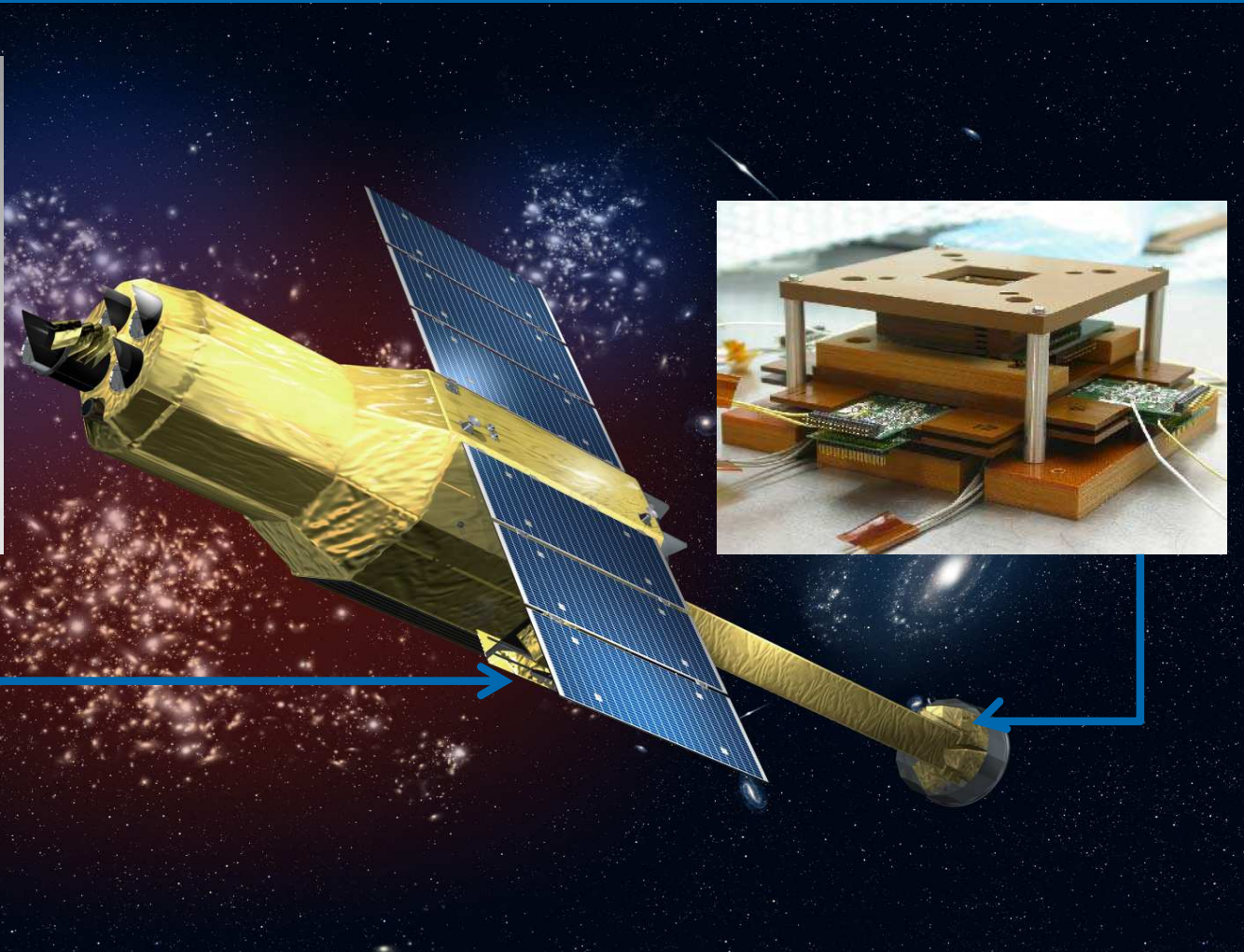
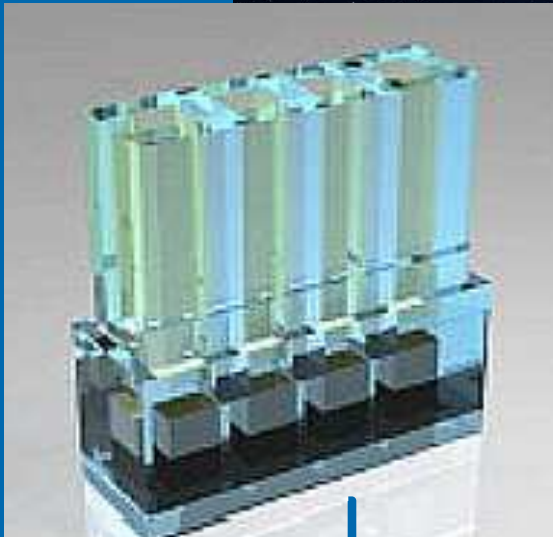
ASTRO-H is the next major X-ray mission with a planned launch date of 2014. Its key scientific objectives include:

- revealing the large scale structure of the Universe and its evolution

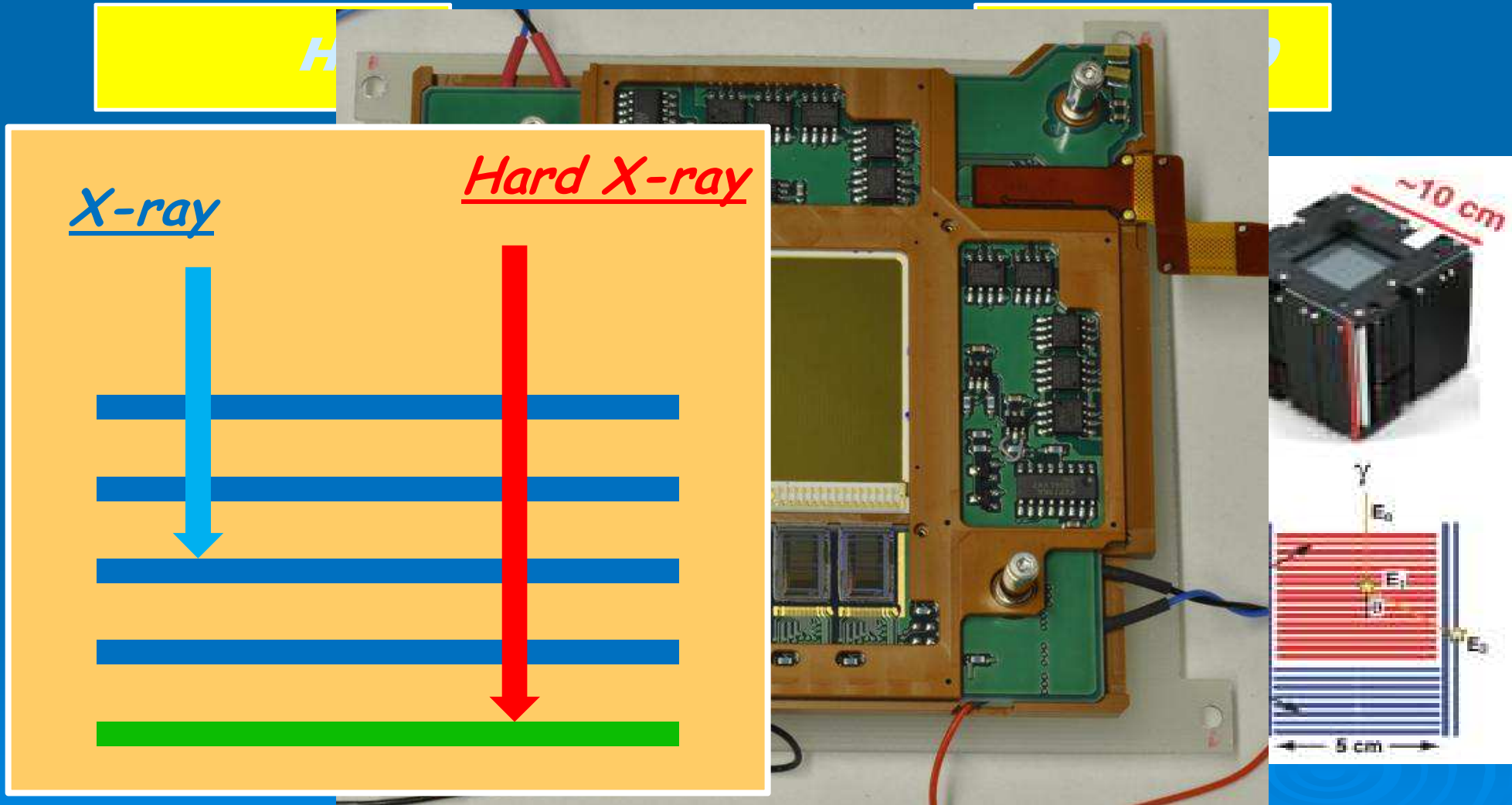
A



ASTRO-H

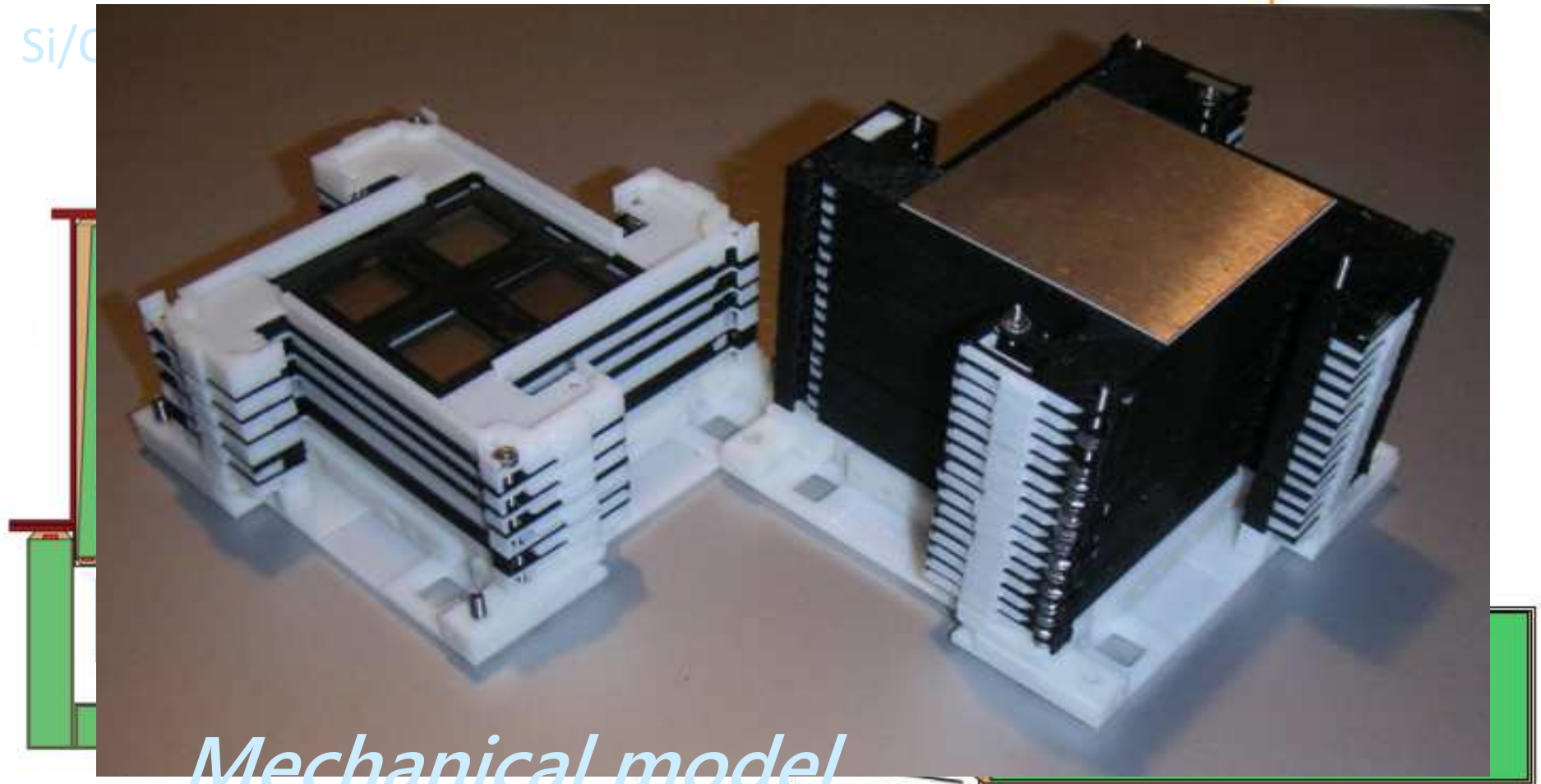


HXI and SGD



HXL and SGD

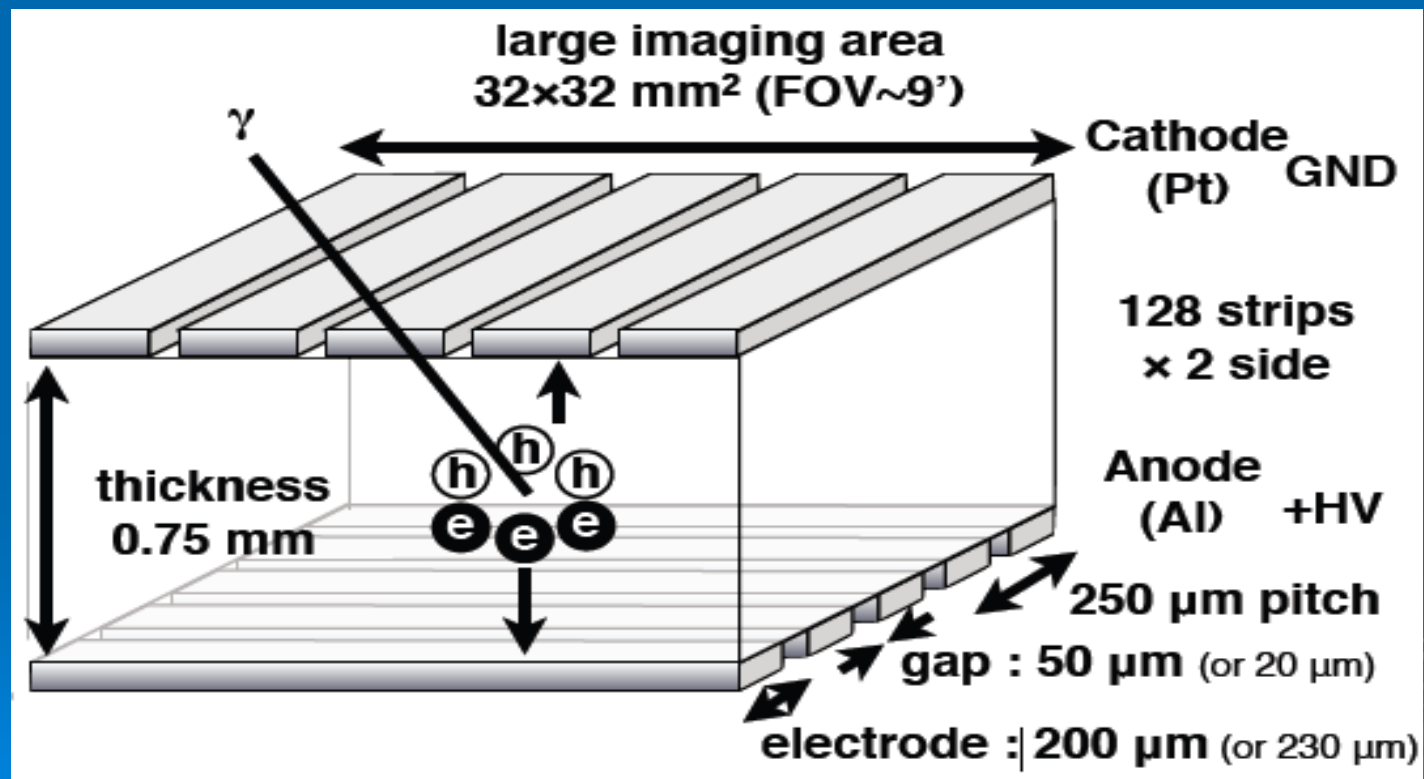
Si/C



*Mechanical model
of SGD*

Detector for HXI

- DS-CdTe : high detection efficiency thanks its atomic number (48-52)

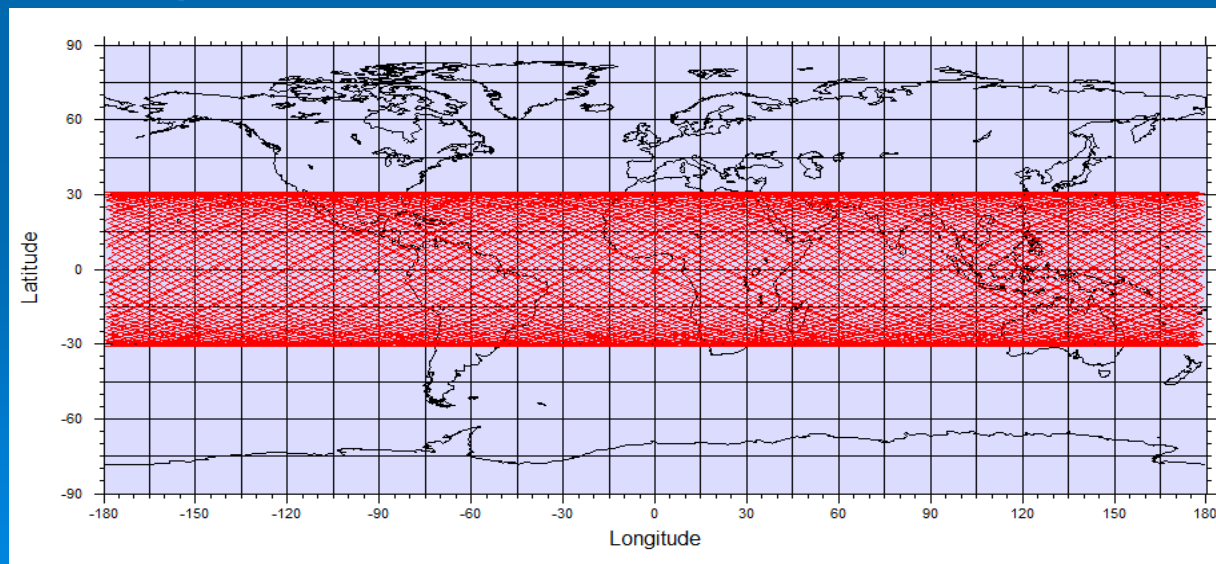


Study of Astro-H SGD ASICs

Astro-H orbit

Low Earth Orbit

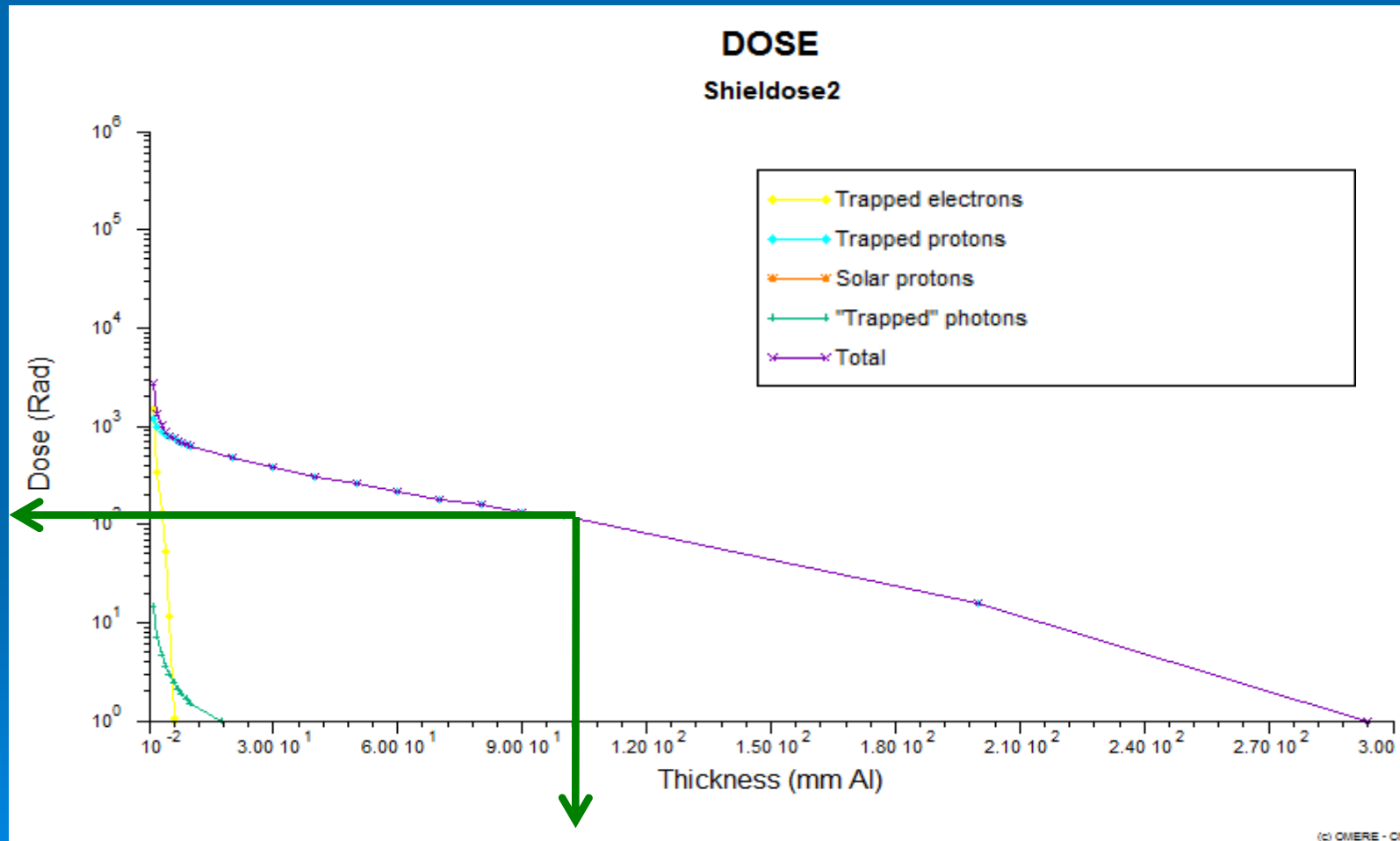
- Orbit Altitude: 550 km
- Orbit Inclination: ~31 degrees
- Orbit Period: 96 minutes
- Launch : 2014
- 4 years of operation



Models used to determine the Astro-H space environment

- a) Trapped protons model is *AP-8 MAX* (NASA/NSSDC)
- b) Trapped electrons model is *AE-8 MIN* (NASA/NSSDC)
- c) Solar protons model is *SOLPRO* (NASA/NSSDC)
- d) Cosmic-ray model is *GCR-ISO* (ISO 15390)

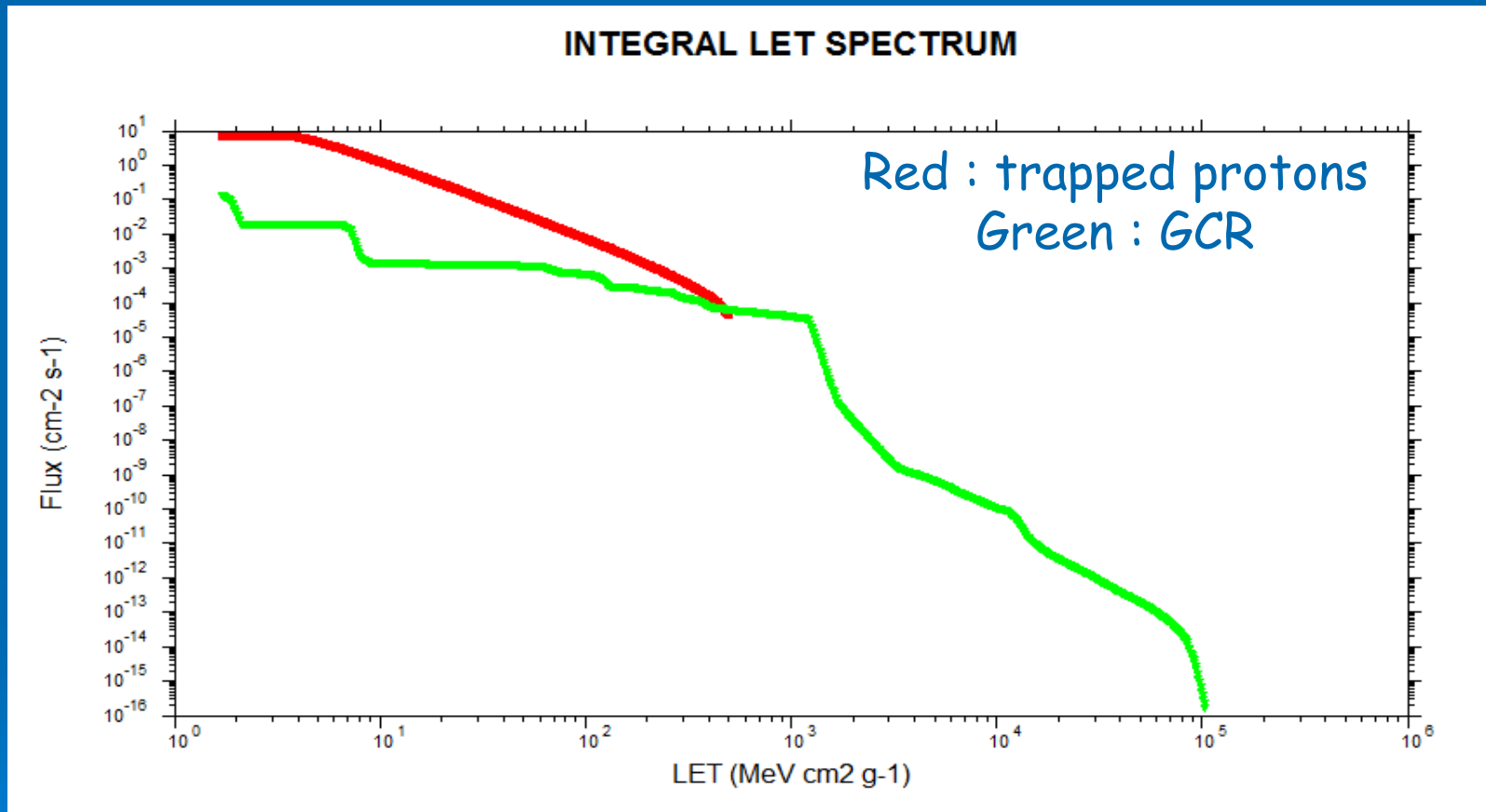
Dose curve



100 rad

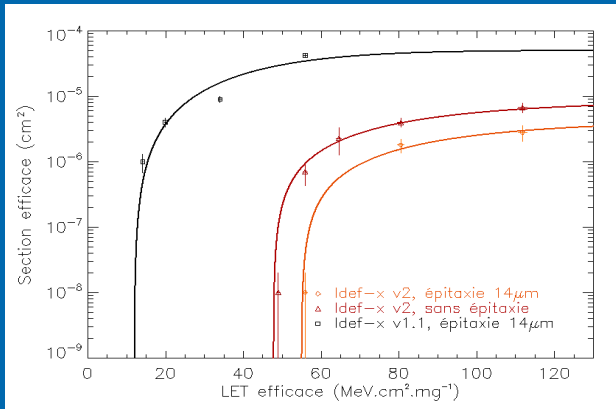
100 mm

LET spectrum



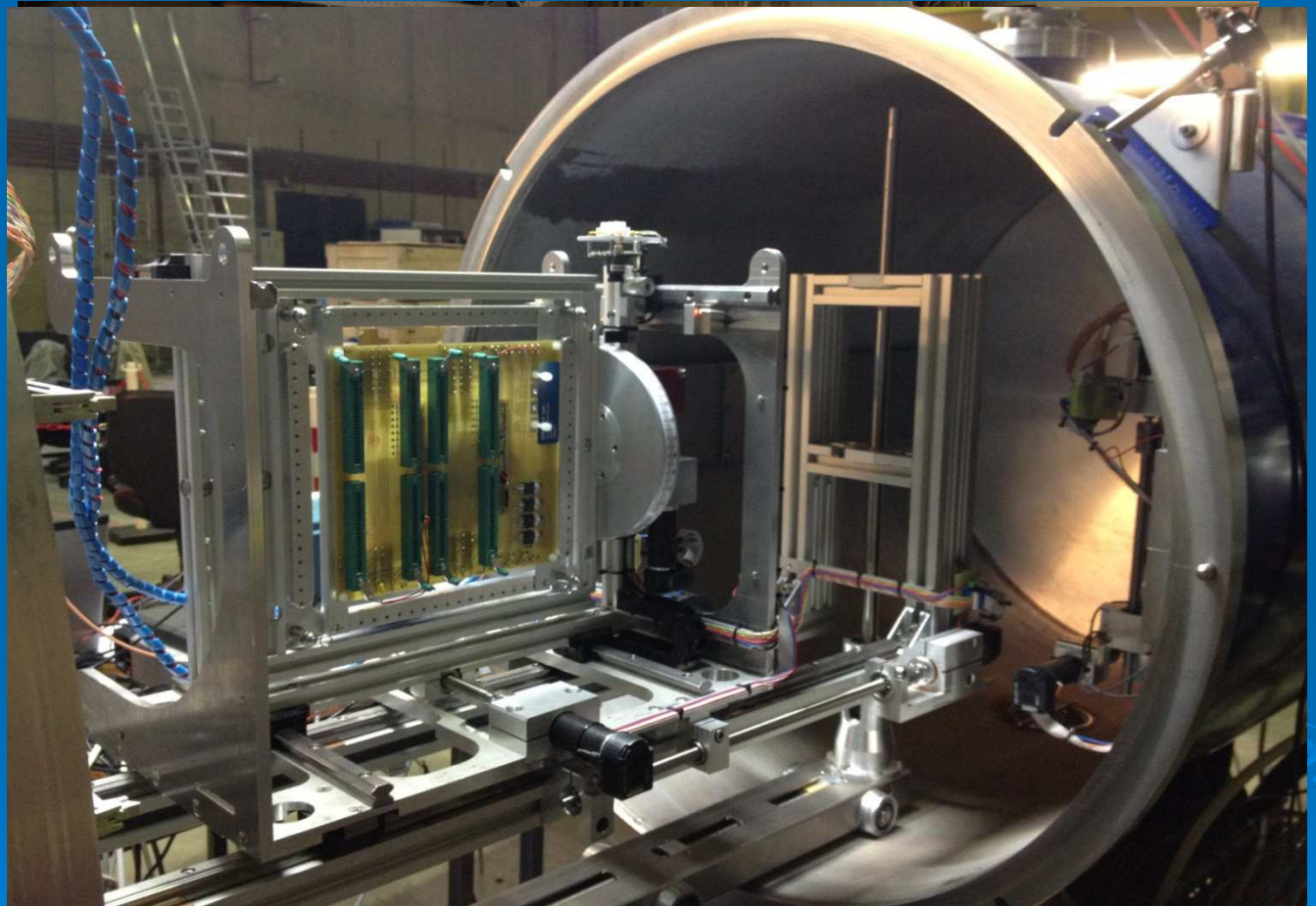
LET computed with 10 cm Al shielding

Radiation Test (1/2)



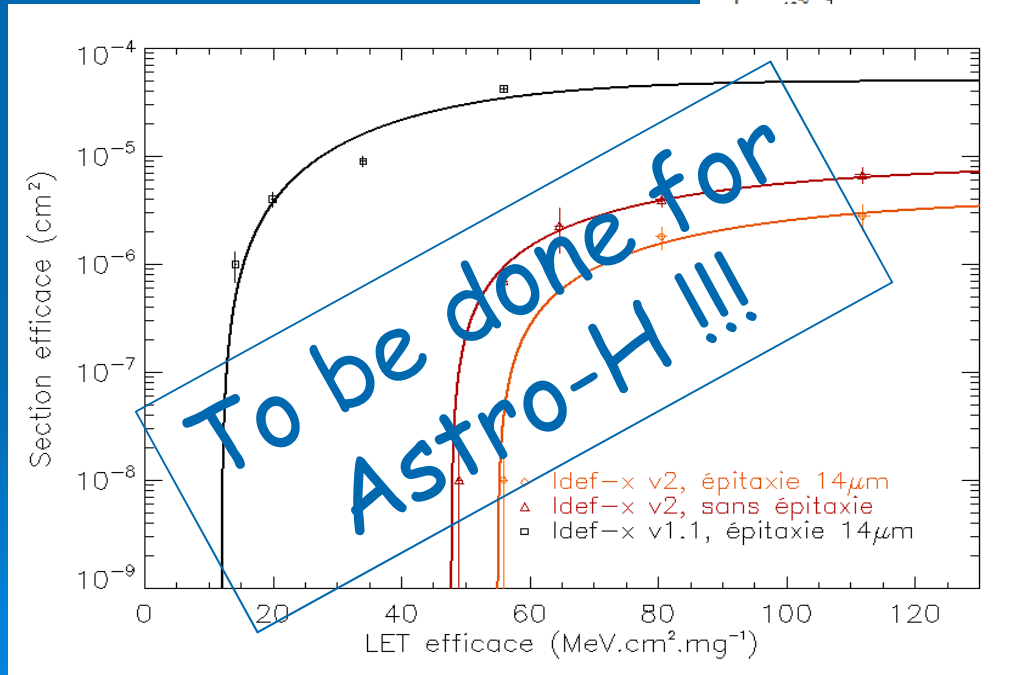
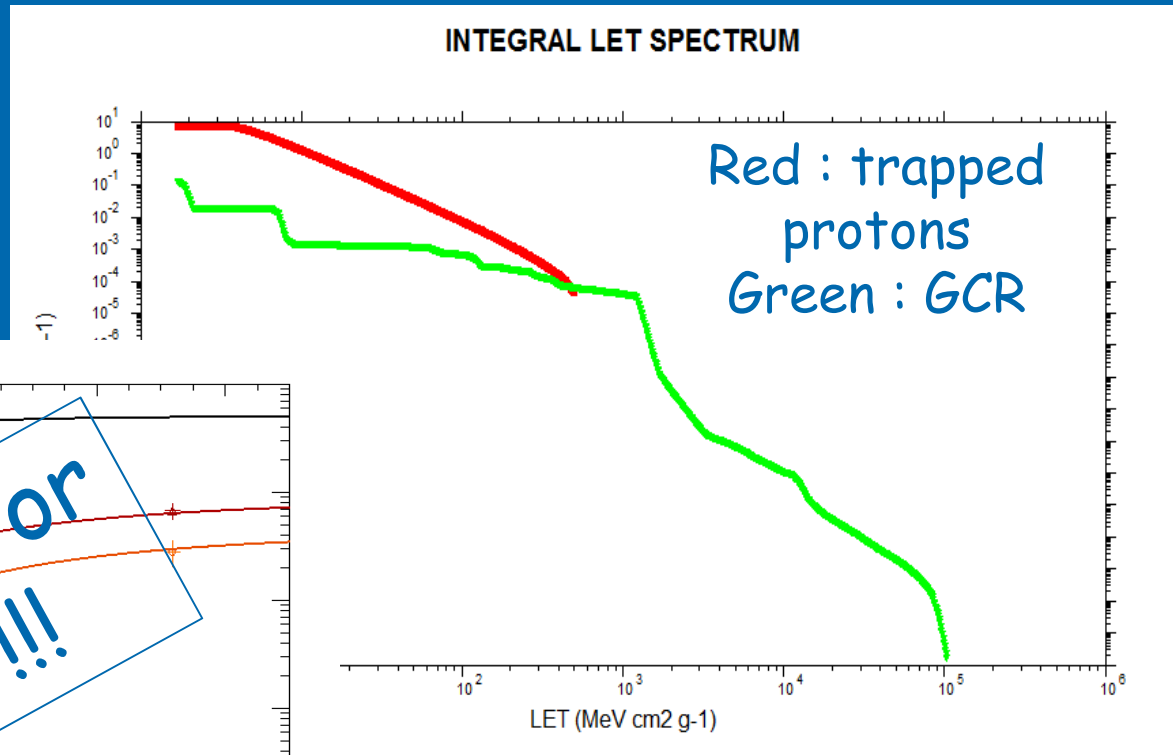
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| ²² Ne ⁺⁷ | 235 | 3.6 | 199 |
| ⁴⁰ Ar ⁺¹² | 372 | 9.95 | 119 |
| ⁵⁸ Ni ⁺¹⁸ | 567 | 21.3 | 98 |
| ⁸³ Kr ⁺²⁵ | 756 | 31.0 | 92 |



HIF accelerator line at Louvain-la-Neuve (Belgium)

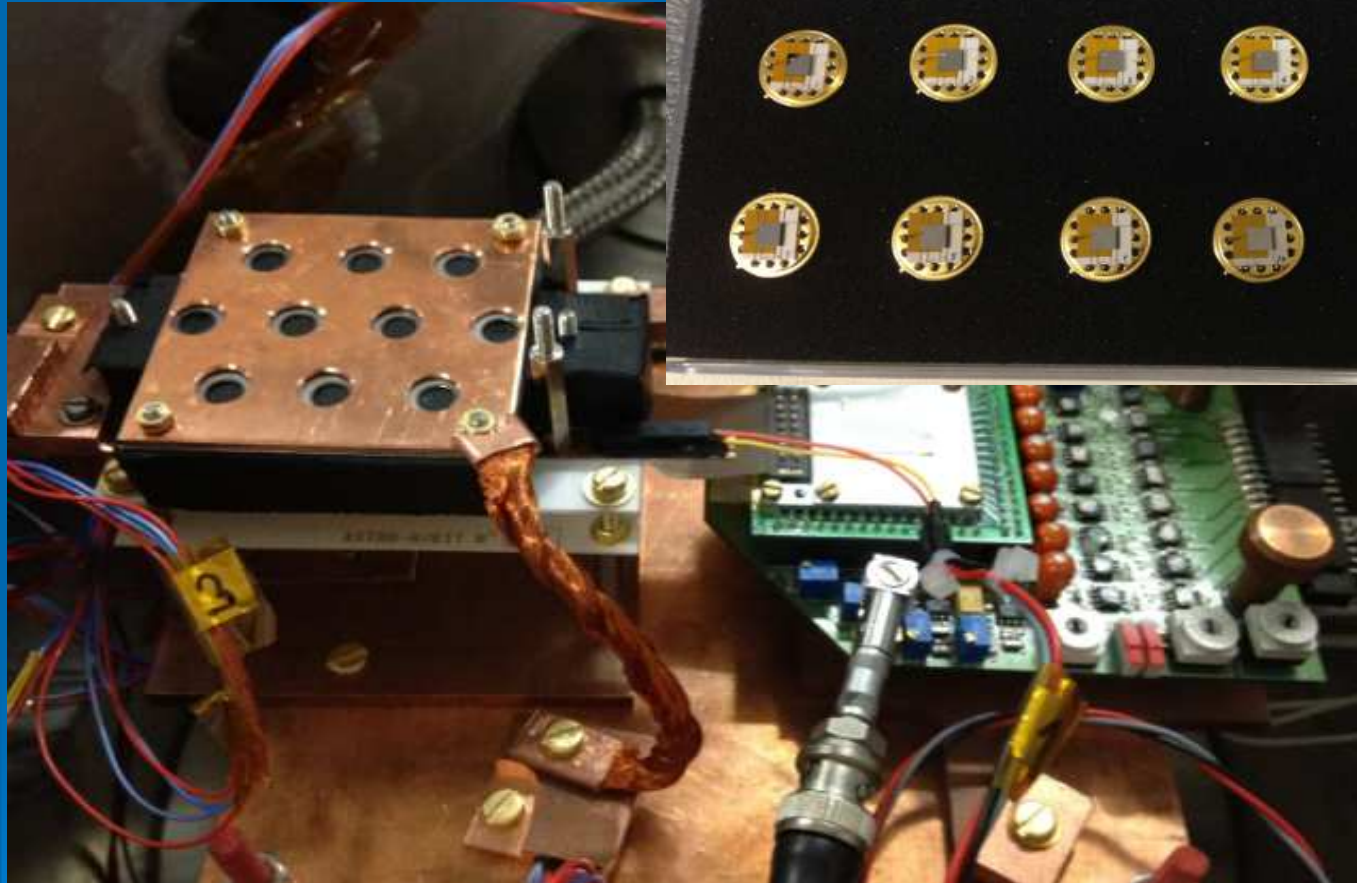
Radiation tests (2/2)



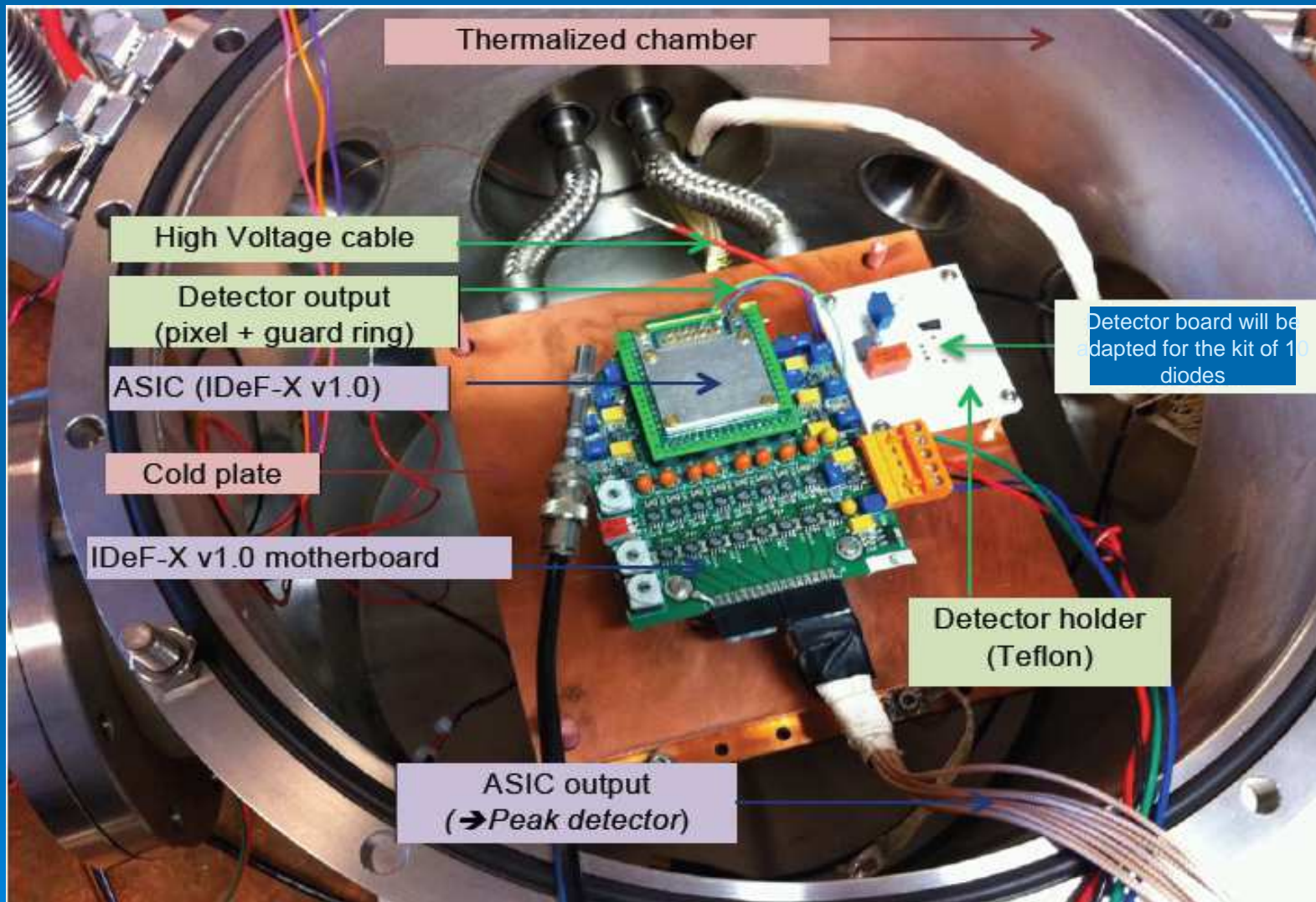
⇒ SEE probability !!

Study of Astro-H HXI and SGD CdTe detectors

Test bench in CEA



test equipment for the diodes



PSI protons accelerator (Switzerland)



Thank you !